

Analysis of extreme rainfall in the Ebre Observatory (Spain)

Núria Pérez-Zanón¹ · M. Carmen Casas-Castillo² · Raúl Rodríguez-Solà³ ·

Juan Carlos Peña⁴ · Anna Rius⁴ · J. Germán Solé⁵ · Ángel Redaño⁶

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Abstract The relationship between maximum rainfall rates for time intervals between 5 min and 24 h has been studied from almost a century (1905–2003) of rainfall data registered in the Ebre Observatory (Tarragona, Spain). Intensity–duration–frequency (IDF) curves and their master equation for every return period in the location have been obtained, as well as the probable maximum precipitation (PMP) for all the considered durations. In particular, the value of the 1-day PMP has resulted to be 415 mm, very similar to previous estimations of this variable for the same location. Extreme rainfall events recorded in this period have been analyzed and classified according to their temporal scale. Besides the three main classes of cases corresponding to the main meteorological scales, local, mesoscale, and synoptic, a fourth group constituted by complex events with high-intensity rates for a large range of durations has been identified also, indicating the

contribution of different scale meteorological processes acting together in the origin of the rainfall. A weighted intensity index taking into account the maximum rainfall rate in representative durations of every meteorological scale has been calculated for every extreme rainfall event in order to reflect their complexity.

1 Introduction 33

Knowledge about spatial distribution and temporal rainfall intensity is essential in meteorology, hydrology, and civil engineering. Rainfall rate behavior in a location has to be considered for planning hydraulic works, roads, and sewage systems and designing rainwater drainage networks in large infrastructures, and evaluate natural risk due to extreme events occurring in the area. As flood is one of the most important natural hazards in mid-latitude Mediterranean areas (Llasat et al. 2013), the main goal of this work is the study of extremely intense rainfall events able to cause flooding and runoff. The intensity–duration–frequency curves (IDF curves) are an important tool for hydrological planning.

Annual mean precipitation registered in the Ebre Observatory (Tarragona, Spain) is around 530 mm, being more than 650 mm in the city of Barcelona. Both are coastal locations of Catalonia (Spain, Fig. 1), a region with a high spatial rainfall variability ranging from less than 400 mm of annual mean precipitation on the western end of the inner area to more than 1250 mm in the highest mountainous zones of the Catalan Pyrenees, and a high seasonal variability (Martínez et al. 2007). Despite the general concordance between the spatial distribution of the mean annual precipitation and daily magnitudes as the maximum daily rainfall or the daily probable maximum precipitation (PMP), remarkable differences were found (Casas et al. 2007, 2008) as a

✉ M. Carmen Casas-Castillo
m.carmen.casas@upc.edu

¹ Center for Climate Change (C3), Campus Terres de l'Ebre, Universitat Rovira i Virgili, Avda. Remolins, 13-15, Tortosa, Spain

² Departament de Física i Enginyeria Nuclear (FEN), EET, Universitat Politècnica de Catalunya, BarcelonaTech (UPC), Campus de Terrassa, Edifici TR2, C/ Colom, 1, 08222 Terrassa, Spain

³ Departament de Física i Enginyeria Nuclear (FEN), EPSEVG, Universitat Politècnica de Catalunya BarcelonaTech (UPC), Víctor Balaguer s/n, 08800 Vilanova i la Geltrú, Spain

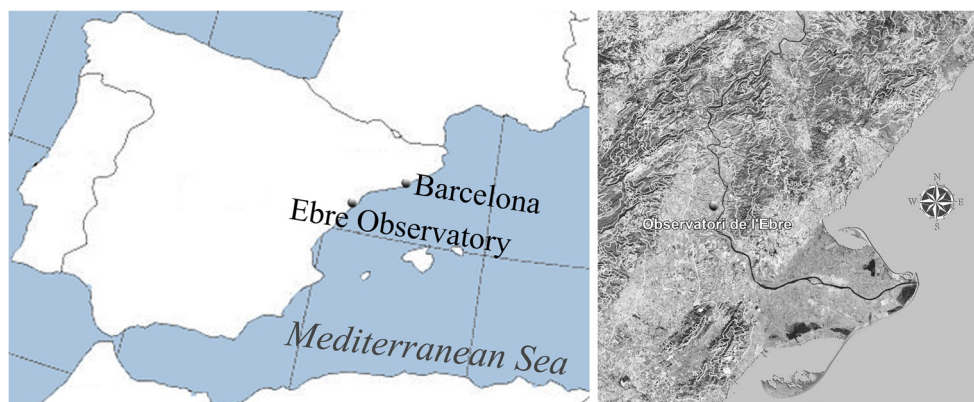
⁴ Servei Meteorològic de Catalunya, Berlin, 38-46, 08029 Barcelona, Spain

⁵ Observatori de l'Ebre, CSIC-URL, Horta Alta, 38, 43520 Roquetes, Spain

⁶ Departament d'Astronomia i Meteorologia (DAM), Facultat de Física, Universitat de Barcelona (UB), Martí i Franqués, 1, 08028 Barcelona, Spain

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Fig. 1 Geographical situation of the Ebre Observatory. At the right, orthophoto of the local area with level curves



59 consequence of the different meteorological situations con- 95
 60 tributing to high rainfall amounts for monthly or annual pe- 96
 61 riods (synoptic scale organizations) and those producing the 97
 62 highest rainfall amounts in 1-day time intervals (local and 98
 63 mesoscale systems). There are several studies dedicated to 99
 64 calculate the IDF curves for Barcelona (Lorente and Redaño 100
 65 1990; Casas et al. 2004; Rodríguez et al. 2014), from the large 101
 66 rainfall data set registered by the Jardí gauge of the Fabra 102
 67 Observatory of the city. In this work, IDF values have been 103
 68 calculated from almost a century of rainfall data (1905–2003) 104
 69 registered in the Ebre Observatory. The relationship between 105
 70 the rainfall intensity values registered for each of the time 106
 71 intervals considered from 5 min to 24 h has been studied for 107
 72 every extreme event recorded in the analyzed period, to inves- 108
 73 tigate the prevailing meteorological temporal scale in each 109
 74 case. The PMP, i.e., “the greatest depth of precipitation for a 110
 75 given duration meteorologically possible for a given size 111
 76 storm area at a particular location at a particular time of year, 112
 77 with no allowance made for long-term climatic trends” 113
 78 (WMO 1986), has been estimated also using a statistical 114
 79 approach. 115

80 **2 Rainfall rate data statistical analysis**

81 The data set used in this study has been obtained from 124
 82 the digitalization of the strip charts recorded between 125
 83 1905 and 2003 by two siphon rain gauges located in 126
 84 the Ebre Observatory. These registers were interrupted 127
 85 due to the Spanish civil war for approximately 1 year, 128
 86 from April 4, 1938, to May 1, 1939, so they last 129
 87 98 years. From these records, the maximum rainfall 130
 88 amounts collected in different time intervals have been 131
 89 obtained. As it is well known, with this type of instru-
 90 ment, it is possible to determine the intensity indirectly
 91 from the record of cumulated rainfall to estimate the
 92 rainfall amounts in any interval of time. Nevertheless,
 93 for short intervals of time as 10 or 20 min, the deter-
 94 mination of the intensity can be affected of a not

negligible error (between 10 and 15 %). The main cause
 of this error is, on one hand, the low temporal resolu-
 tion of the instrument (10 min corresponds approximate-
 ly to 3 mm of band) and, on the other hand, that the
 instrument takes between 10 and 20 s to download the
 accumulated water whenever it gathers the equivalent of
 10 mm of rainfall (Lanza et al. 2005). In spite of this, it
 is not usual to have series of cumulated rainfall so ex-
 tensive, so complete, and with such verifiable quality
 like those obtained from this observatory, which can
 contribute valuable information to the knowledge of
 the climatology of this variable in this zone. Thus, the
 maximum rainfall amounts collected in different time
 intervals between 5 min and 24 h (5, 10, 15, 20, 25,
 30, 35, 40, 45, 50, and 55 min, and 1, 2, 6, 12, and
 24 h) have been calculated. For each of the 16 consid-
 ered durations, the two highest values per year have
 been selected (i.e., a peaks-over-threshold approach con-
 sidering a return period of 0.5 years as threshold), lead-
 ing to 16 data series of 196 values each. These series
 have been considered stationary since no significant
 trends (p value <0.05) have been detected by a Mann–
 Kendall test (Yue et al. 2002). In order to analyze their
 statistical distribution, several theoretic functions have
 been tested. Among them, the generalized Pareto distri-
 bution (GPA) has provided the best fit for the data. The
 cumulative distribution function of the GPA is defined
 by Eq. (1),

$$F(x) = 1 - \left[1 - \kappa \left(\frac{x - \varepsilon}{\alpha} \right) \right]^{\frac{1}{\kappa}} \quad (1)$$

where x is the maximum rainfall amount, ε the location
 parameter, α the scale parameter, and κ the shape pa-
 parameter of the GPA distribution. These parameters have
 been estimated using the L-moments method (Hosking
 and Wallis 1997). Figure 2 shows the L-moments dia-
 gram with the theoretical curve of the GPA distribution
 fitted to the L-moments values (circles) calculated from
 the data series. The estimated parameter values for each

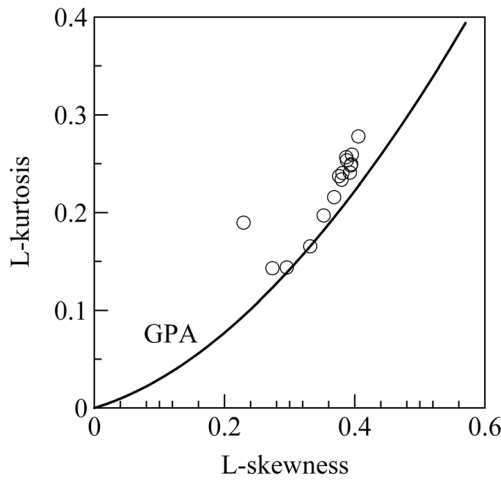


Fig. 2 L-moments diagram calculated for the rainfall data series recorded in the Ebre Observatory (1905–2003). GPA distribution in the *solid line*

132 of the 16 considered durations are listed in Table 1.

133 **3 Determination of the IDF curves**

134 From the fitted GPA distributions, the estimated maximum
 135 rainfall amounts for every duration t and 1-, 2-, 5-, 10-, 25-,
 136 50-, and 100-year return periods T have been determined
 137 (Table 2), and so the intensity values $I(t, T)$ shown in Fig. 3.
 138 The relationships between these magnitudes are known as IDF
 139 curves. Among the different empirical fitting functions used
 140 for IDF curves, Sherman's equation (Eq. (2)) (Chen 1983) has
 141 been considered to fit the $I(t)$ values for every T ,

$$I(t) = \frac{a}{(b + t)^c} \tag{2}$$

142 where a , b , and c are fitting parameters. Following a method
 143 proposed by Chen (1983), these $I(t)$ values have been rescaled
 144 using the intensity $I(60, T)$ as rescaling factor to determine a
 145 master equation (Casas et al. 2004) relating the three variables
 146 $I(t, T)$. The ratio $\frac{I(t, T)}{I(60, T)}$ between the maximum intensity values
 147 for each return period and every duration, $I(t, T)$, and the 1-h
 148 maximum intensity for the same return period, $I(60, T)$, has
 149 been fitted using the Sherman equation (Eq. (2)). The fitting
 150 procedure has been done determining first the value of the

Table 2 Estimated maximum rainfall amount (in mm) for every duration t and return period T

t (min)	T (years)								
	1	2	5	10	15	25	50	100	
5	8.5	10.0	12.0	13.4	14.2	15.3	16.7	18.1	t2.4
10	12.8	15.6	19.0	21.3	22.6	24.1	26.1	27.8	t2.5
15	15.9	19.5	23.9	27.1	28.9	31.0	33.8	36.4	t2.6
20	18.1	22.3	27.7	31.7	34.0	36.8	40.5	44.2	t2.7
25	19.7	24.5	30.9	35.7	38.4	41.9	46.5	51.1	t2.8
30	21.0	26.4	33.5	38.9	42.1	46.2	51.7	57.3	t2.9
35	22.2	28.0	35.9	42.0	45.7	50.3	56.9	63.5	t2.10
40	23.3	29.5	38.2	45.0	49.1	54.5	62.0	69.8	t2.11
45	24.3	30.9	40.2	47.6	52.1	58.0	66.2	74.9	t2.12
50	25.2	32.2	42.0	49.9	54.7	61.0	70.0	79.5	t2.13
55	26.0	33.2	43.5	51.9	57.0	63.7	73.3	83.6	t2.14
60	26.7	34.2	44.8	53.5	58.8	65.8	75.8	86.5	t2.15
120	32.7	42.4	56.5	68.2	75.5	85.2	99.3	114.6	t2.16
360	42.7	55.2	73.4	88.5	98.0	110.6	129.0	149.1	t2.17
720	50.6	64.7	85.6	103.3	114.5	129.5	151.8	176.3	t2.18
1440	58.8	76.5	102.6	124.4	138.1	156.4	183.3	212.6	t2.19

parameter c which can be related to the temporal scaling prop-
 152 erties of the rainfall series (Rodríguez et al. 2014). It has been
 153 widely observed that the probability distribution of the annual
 154 maximum precipitation intensities satisfies scale relationships
 155 (Koutsoyiannis and Foufoula-Georgiu 1993; Burlando and
 156 Rosso 1996; Menabde et al. 1999), which means that the
 157 probability distribution of the annual maximum intensity for
 158 a given duration, i.e., $t_0=60$ min, $I(60)$, and the distribution at
 159 the other scale $I(t)$ (duration $t=\lambda t_0$), will be related by a factor
 160 that is a power function of the scale parameter λ , $I(t)=\lambda^\beta I(t_0)$,
 161 β being a scaling exponent. In terms of the IDF values, the
 162 scaling relationship leads to expression (3).
 163

$$I(t, T) = I(60, T) \left(\frac{t}{60}\right)^\beta \tag{3}$$

Comparing Eq. (3) and Sherman's equation (Eq. (2)), ex-
 164 ponent c can be identified as $-\beta$ if $b=0$. Casas et al. (2004) and
 165 Rodríguez et al. (2014) found a value of b around 10 min for
 166 Barcelona, which can be considered small comparing to t for
 170

t1.1 **Table 1** GPA distribution parameters ε (location), α (scale), and κ (shape) estimated using the L-moments method, for each of the 16 considered durations t

t1.2 t (min)	5	10	15	20	25	30	35	40	45	50	55	60	120	360	720	1440
t1.3 ε	7.01	9.72	12.09	13.7	14.74	15.75	16.62	17.27	17.92	18.61	19.2	19.67	23.81	31.27	37.8	42.51
t1.4 α	2.19	4.6	5.62	6.35	7.13	7.59	8	8.49	8.95	9.26	9.51	9.81	12.36	15.84	17.54	22.43
t1.5 κ	0.018	0.118	0.079	0.038	0.014	-0.013	-0.038	-0.057	-0.067	-0.079	-0.089	-0.091	-0.117	-0.121	-0.142	-0.128

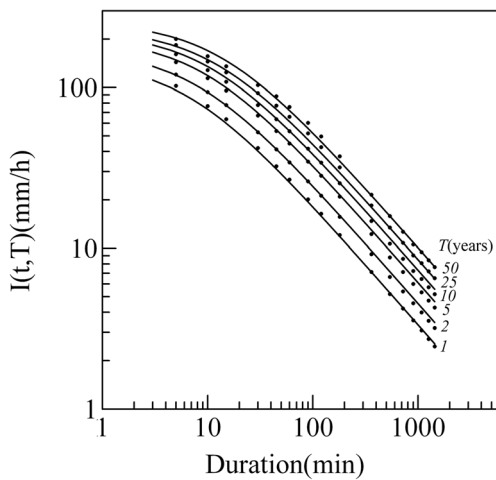


Fig. 3 Intensity values $I(t, T)$ (in mm/h) for durations t between 5 min and 24 h and return periods from 1 to 50 years, and IDF curves

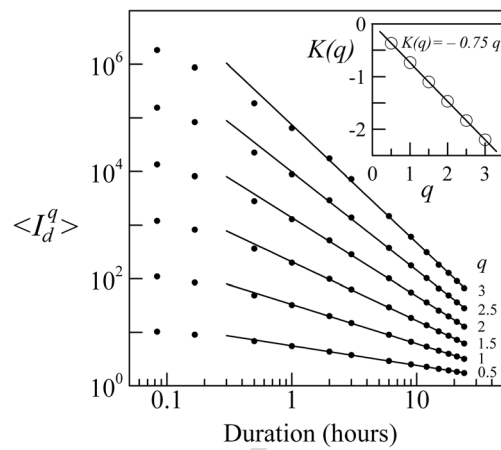


Fig. 4 Statistical moments for different values of q of the annual maximum intensity of the Ebre series, as a function of the duration (between 5 min and 24 h). The fitted straight lines indicate scale invariance over the range from 1 to 24 h. Their slopes determine the scale function $K(q)$. The inserted figure shows a linear regression between $K(q)$ and the order q of the moments

171 durations larger than 1 h. This scaling exponent β has been
 172 calculated for the annual maximum rainfall series registered in
 173 the Ebre Observatory using their statistical moments as pro-
 174 posed by Menabde et al. (1999). The scaling relationship be-
 175 tween distributions can be expressed in terms of their mo-
 176 ments of order q , $\langle I_d^q \rangle = \lambda^{K(q)} \langle I_{d_0}^q \rangle$. If the scale function
 177 $K(q)$ is linear, $K(q) = \beta q$, the process is a simple scale process
 178 or monofractal. If not, the process is multiscale or
 179 multifractal. A simple scaling regime between 1 and 24 h with
 180 a value of $\beta = -0.75$ has been found (Fig. 4). Then, the ratio
 181 $\frac{I(t, T)}{I(60, T)}$ has then been fitted to Sherman's equation (Eq. (2))
 182 fixing parameter c as 0.75 (± 0.1). The obtained parameters a
 183 and b have been listed in Table 3. Whereas parameter a result-
 184 ed very similar for every return period (mean value of 23.9,
 185 close to $(1/60)^{-0.75}$, with a standard deviation of 0.5), param-
 186 eter b is slightly different for every return period (between 6
 187 and 16 min; see Table 3). The behavior of parameter b with
 188 return period T can be described by a power expression as $b =$
 189 $6.9T^{0.19}$ (6.9 ± 0.2 , 0.19 ± 0.01). Considering this, the fitting by
 190 Sherman's equation has been found to be Eq. (4),

$$\frac{I(t, T)}{I(60, T)} = \frac{23.9}{(6.9T^{0.19} + t)^{0.75}} \quad (4)$$

192 where $I(t, T)$ is the maximum intensity in millimeters per min-
 193 ute for a duration of t minutes with a T -year return period. The
 194 intensity of reference $I(60, T)$ shows in turn a logarithmic de-
 195 pendence with the return period T (Eq. 5),

$$I(60, T) = 0.496 \log T + 0.42 \quad (5)$$

196 where T is expressed in years and the intensity $I(60, T)$ in
 197 millimeters per minute. Combining Eqs. (4) and (5), the mas-
 198 ter equation for the IDF curves of the Ebre Observatory has
 199 been obtained (Eq. 6).
 200

$$I(t, T) = \frac{11.9 \log T + 9.9}{(6.9T^{0.19} + t)^{0.75}} \quad (6)$$

The parameters of the numerator in Eq. (6) have been found to be 11.9 ± 0.6 and 9.9 ± 0.6 .

This master equation fits satisfactorily the IDF points of Table 2 with a relative error lower than 5 % for durations between 5 min and 6 h and return periods from 1 to 50 years. For the remaining cases, the error is lower than 10 %, as expected considering the uncertainties found for every parameter of Eq. (6). The corresponding curves to the 1-, 2-, 5-, 10-, 25-, and 50-year return periods have been shown in Fig. 3.

4 Classification of extreme rainfall events

From the Ebre Observatory records, rainfall events showing intensities equal to or larger than the 10-year return period intensity values given by Eq. (6) for any duration have been selected. The considered durations have been the 16 time

Table 3 Fitting parameters a and b for every return period T

T (years)	a	b	t3.1 t3.2
1	23.1	6.6	t3.3
2	23.5	7.9	t3.4
5	23.7	9.7	t3.5
10	23.9	10.7	t3.6
15	23.9	11.8	t3.7
25	24.1	12.8	t3.8
50	23.9	13.9	t3.9
100	24.7	16.1	t3.10

218 intervals listed in Tables 2 and 5 more (180, 540, 900,
 219 1080, and 1260 min) added to complete the rainfall infor-
 220 mation between 2 and 24 h in order to obtain a better
 221 representation of this temporal period. In the analyzed
 222 98-year period, 28 rainfall events have reached or
 223 exceeded these intensities in one or more of the studied
 224 durations, 71 % of them occurring in autumn. The rela-
 225 tionship between the recorded rainfall amounts for every
 226 time interval may provide useful information on the origin
 227 and nature of the rainfall. Cluster analysis (Anderberg
 228 1973) is an appropriate technique to investigate this rela-
 229 tionship and classify the extreme events based on their
 230 similarity (Romero et al. 1999; Casas et al. 2004, 2010).
 231 Every rainfall event i is characterized by n measured or
 232 calculated variables ($x_i^1, x_i^2, \dots, x_i^n$). These variables can
 233 be considered as coordinates representing the i rainfall
 234 event in an n -dimensional space, so the similarity concept
 235 between the events can be specified based on the
 236 Euclidean distance between the corresponding points in
 237 this n -dimensional space. The process leading to the clus-
 238 ters or groups selection is agglomerative. Among the dif-
 239 ferent clustering procedures, the UPGMA (unweighted
 240 pair group method using arithmetic averages), one of the
 241 most widely used, has been chosen. The main reason of
 242 this choice is that the use of this aggregation method
 243 resulted in the better adapted dendrogram to our aim of
 244 relating the obtained clusters to the temporal scales of the
 245 meteorological processes that originate the storms. The
 246 UPGMA usually provides clusters neither too big nor
 247 small, with similar and small variances, and not too sen-
 248 sitive to the presence of outliers in the sample. Using
 249 UPGMA, the distance between two clusters is defined as
 250 the averaged distance between every pair formed by an
 251 event from one cluster and another from the other one.
 252 The variables used to characterize every one of the select-
 253 ed 28 rainfall events are the 21 recorded maximum rain-
 254 fall rates within the analyzed time intervals. The obtained
 255 dendrogram is shown in Fig. 5, where the nearest (so
 256 more similar) events are those connected by the shortest
 257 path along the branches. The first vertical line drawn
 258 (grey line L1) intersecting the hierarchical tree divides
 259 one group of events (group IV in Fig. 5) from the rest
 260 of the cases. This group is constituted by events 181040,
 261 191065, 260992, and 150943 (ddmmyy format), showing
 262 a high rainfall intensity ($T > 10$ years) for a wide range of
 263 durations. Three of them exceed the imposed threshold
 264 between 10 min and 24 h, and the remaining one does
 265 for the whole considered temporal range (5 min–24 h).
 266 These rainfall events, which caused rain flood in the area,
 267 are the most complex storms and show the joint action of
 268 processes corresponding to more than one of the
 269 meteorological scales, as well as the events classified in
 270 group IV by Casas et al. (2004, 2010) for extreme storms

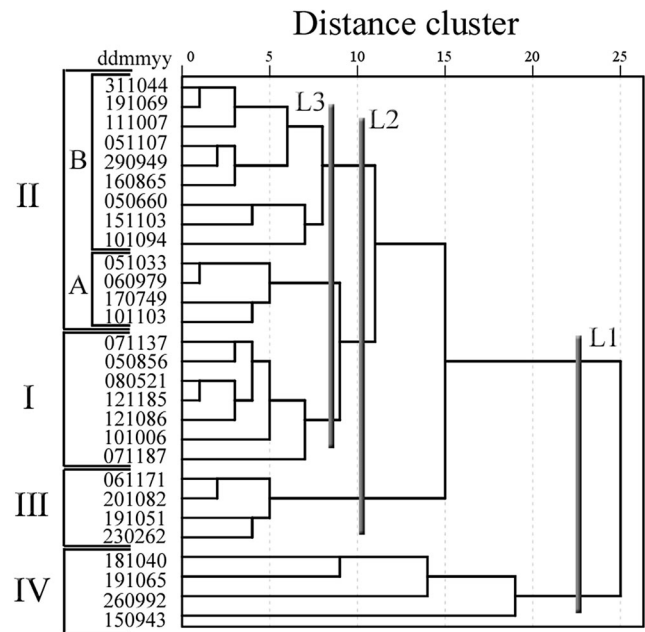
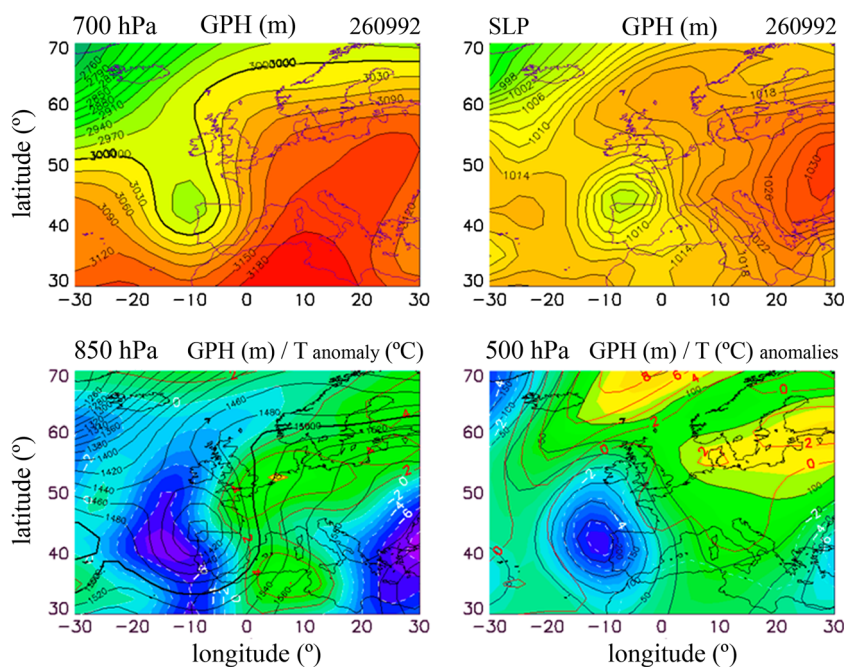


Fig. 5 Dendrogram of the 28 selected extreme rainfall events (ddmmyy)

271 recorded in Barcelona. The synoptic situations corre- 271
 272 sponding to this kind of complex events (14 % of the 272
 273 studied cases) are characterized by the presence of a 273
 274 depression located SW of the Iberian Peninsula and a 274
 275 powerful blocking anticyclone on the European continent. 275
 276 In some cases (Fig. 6), the depression is clearly reflected 276
 277 on the surface, as happened on event 260992, whereas in 277
 278 others, it only appears in height being much weaker or 278
 279 nonexistent on the surface (Fig. 7). In these situations, 279
 280 the prevailing wind on Catalonia is from the S or SW in 280
 281 medium and high levels, while at 850 hPa and on the 281
 282 surface, the wind has a strong component of E, resulting 282
 283 in a moist and warm advection from the Mediterranean 283
 284 Sea. These situations produce long-lasting advective pre- 284
 285 cipitation (large scale) at the same time that convective 285
 286 phenomena linked to the SE warm flow in low levels, as 286
 287 mesoscale convective systems with even embedded small- 287
 288 er convective cells, cause extreme intense rainfall. 288
 289 Drawing a second vertical line (grey line L2, Fig. 5), 289
 290 three more groups can be identified. Four events (group 290
 291 III in Fig. 5) have a clear synoptic origin. The other two 291
 292 groups are group IIB and a large ensemble that can be 292
 293 divided using a third vertical line (grey line L3) into 293
 294 groups I and IIA. Group I is formed by seven events 294
 295 reaching high rainfall rate values (return period equal to 295
 296 or larger than 10 years) only for durations equal to or 296
 297 shorter than 15 min and lower values for the rest, 297
 298 whereas the four events of group IIA show high rates 298
 299 till 35–60 min. Group I is representative of extremely 299
 300 local or microscale rainfall events, with a clear seasonal 300
 301 influence (seven of the eight cases were recorded 301

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Fig. 6 Synoptic situation corresponding to the event 260992 at the sea surface level (sea level pressure (SLP)) and 850, 700, and 500 hPa. The geopotential height and its anomaly at 500 hPa are measured in meters. The temperature anomaly at 850 and 500 hPa is measured in degrees Celsius. Data obtained from the 20th Century V2 Reanalysis Project (20CRP) provided by NOAA/OAR/ESRL PSD, Boulder, CO (Compo et al. 2011)



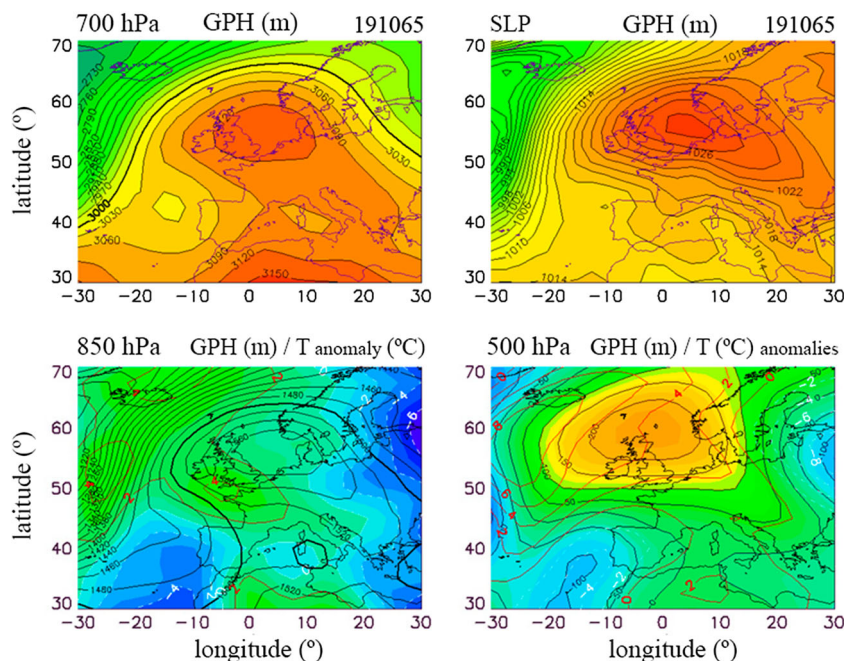
302 November) and an evident diurnal surface heating effect
 303 in the convective development (almost all cases took
 304 place after midday). Group IIA can be identified as
 305 meso- γ -scale convective systems, at the edge of micro-
 306 scale, whereas group IIB events, with high intensities
 307 even for 12–24 h, are compatible with meso- α and
 308 meso- β scales (Thunis and Bornstein 1996). The meteo-
 309 rological situations causing this type of rainfall in the
 310 Mediterranean area are often very active fronts moving

slowly with intense mesoscale rainfall systems developed
 inside (Houze and Hobbs 1982; Browning 1990), or me-
 soscale convective complexes (Codina et al. 1997).

5 Rainfall intensity weighted index

Casas et al. (2004) proposed a rainfall intensity weighted in-
 dex (IP) to reflect the severity and complexity of the registered

Fig. 7 Synoptic situation corresponding to the event 191065 at the sea surface level (SLP) and 850, 700, and 500 hPa. The geopotential height and its anomaly at 500 hPa are measured in meters. The temperature anomaly at 850 and 500 hPa is measured in degrees Celsius. Data obtained from the 20th Century V2 Reanalysis Project (20CRP) provided by NOAA/OAR/ESRL PSD, Boulder, CO (Compo et al. 2011)



317 rainfall events according to the contribution of the different
 318 meteorological scales to the origin of rainfall (Pinto et al.
 319 2013). For each storm, the maximum intensities I_5 , I_{60} , I_{120} ,
 320 and I_{1440} , corresponding to 5-min and 1-, 2-, and 24-h dura-
 321 tions, respectively, are selected as rainfall rate indexes to rep-
 322 resent the local scale, small mesoscale, large mesoscale, and
 323 synoptic scale contributions. The intensity weighted index IP
 324 is defined by Eq. 7,

$$IP(T) = 1/4 \left\{ \frac{I_5}{I(5, T)} + \frac{I_{60}}{I(60, T)} + \frac{I_{120}}{I(120, T)} + \frac{I_{1440}}{I(1440, T)} \right\} \quad (7)$$

326 where T is the chosen return period for the index normaliza-
 327 tion. The highest values of the IP are reached usually in storms
 328 presenting high rainfall intensities for each of the time inter-
 329 vals corresponding to every meteorological scale, even though
 330 a singular storm could have a high IP index for the extraordi-
 331 nary contribution of an only meteorological scale. This index
 332 can be useful to provide complementary information to the
 333 cluster analysis performed in the former section quantifying
 334 the degree of exceptionality of each individual storm.
 335 Choosing T as 10 years, the IP for the 28 extreme rainfall
 336 events has been calculated (Fig. 8). As expected, the highest
 337 indexes correspond to the complex rainfall events classified
 338 into group IV, with values higher than 1.26, and two events
 339 classified as group IIB by the cluster analysis have an IP value
 340 higher than 1. It is remarkable that 68 % of the studied rainfall
 341 events only achieved the imposed intensity level in one or
 342 none of the selected representative durations, seven of the nine
 343 remaining cases show 10-year return period intensities in two
 344 or three time intervals, and only two of the events have
 345 reached them in the four durations simultaneously.

346 6 Maximum probable precipitation estimation

347 Hydrologists use the PMP magnitude and its spatial and tem-
 348 poral distributions to calculate the probable maximum flood
 349 (PMF), which is one of several conceptual flood events used
 350 in the design of hydrological structures, for maximum reliabil-
 351 ity and safety. Procedures for determining the PMP are admit-
 352 tedly inexact: results are estimates and a risk statement has to
 353 be assigned to them. Quoting Koutsoyiannis (1999), the PMP
 354 approach “by no means implies zero risk in reality.” The
 355 National Research Council (1994) estimates the return period
 356 of the PMP in the USA as between 10^5 and 10^9 years. To
 357 estimate the PMP in a place, a variety of procedures based
 358 on the location of the project basin, availability of data, and
 359 other considerations have been proposed (e.g., Wiesner 1970;
 360 Schreiner and Reidel 1978; WMO 1986; Collier and Hardaker
 361 1996). Most of them are based on meteorological analysis,
 362 while some are based on statistical analysis. Among the

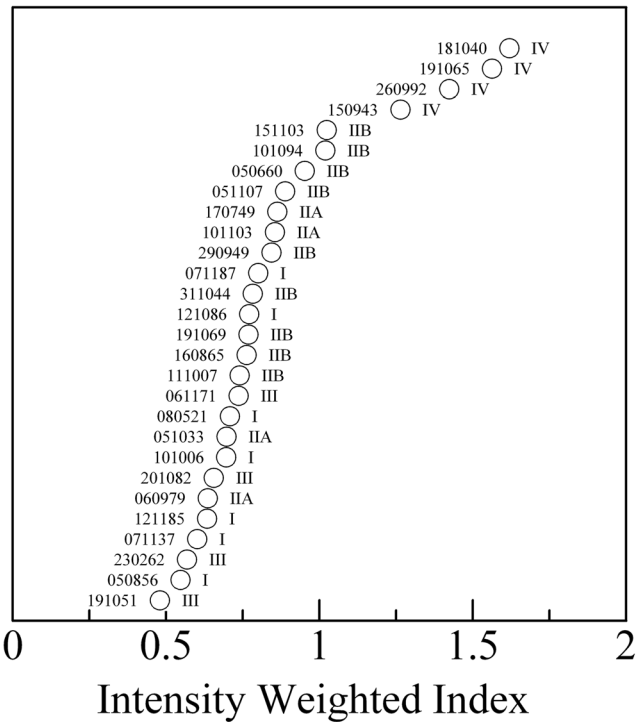


Fig. 8 IP index showing the severity and complexity of the classified extreme rainfall events

statistical methods for estimating the PMP, the most widely 363
 used is the method of Hershfield (1961, 1965), based on the 364
 frequency analysis of the annual maximum rainfall data reg- 365
 istered at the site of interest. In order to estimate the PMP 366
 in the area of study, for durations from 5 min to 24 h, the 367
 Hershfield statistical method has been applied to rainfall data 368
 from the Ebre Observatory (1905–2003). 369

Table 4 Mean \bar{x}_n , coefficient of variation CV, maximum value x_M , and 363
 observed frequency factor k_m of the Ebre Observatory annual maximum 364
 rainfall series for durations t from 5 min to 24 h 365

t (min)	\bar{x}_n (mm)	CV	x_M (mm)	k_m	t4.2
5	9.2	0.3	19.6	3.5	t4.3
10	14.2	0.4	30.5	3.3	t4.4
30	24.5	0.4	68.3	4.8	t4.5
60	32.3	0.5	98.9	4.7	t4.6
120	40.4	0.5	138.6	5.2	t4.7
180	44.8	0.5	166.9	6.0	t4.8
360	54.0	0.5	176.4	5.0	t4.9
540	60.0	0.5	218.6	6.3	t4.10
720	64.8	0.5	228.5	6.1	t4.11
900	68.1	0.5	235.9	5.9	t4.12
1080	71.7	0.5	243.5	5.4	t4.13
1260	74.2	0.5	250.6	5.3	t4.14
1440	76.4	0.5	257.6	5.2	t4.15

t5.1 **Table 5** PMP values in the Ebre
t5.2 Observatory for every considered
duration t in minutes

t (min)	5	10	30	60	120	180	360	540	720	900	1080	1260	1440
PMP (mm)	36	59	112	168	224	250	290	319	338	354	384	400	416
Error (mm)	3	5	10	15	20	23	26	29	30	32	35	36	38

370 The Hershfield technique for estimating the PMP is based
371 on Chow's (1951) general frequency Eqs. (8) and (9),

374
$$\text{PMP} = \bar{x}_n + k_m \sigma_n \tag{8}$$

373
$$k_m = \frac{x_M - \bar{x}_{n-1}}{\sigma_{n-1}} \tag{9}$$

376 where x_M , \bar{x}_n , and σ_n are the highest, mean, and standard
377 deviation for a series of n annual maximum rainfall values
378 of a given duration, \bar{x}_{n-1} and σ_{n-1} are respectively the mean
379 and standard deviation for this series excluding the highest
380 value x_M , and k_m is a frequency factor. Since the frequency
381 factor is the number of standard deviations σ_{n-1} to be added to
382 the mean \bar{x}_{n-1} to achieve the maximum x_M , its value results
383 higher for series including an extraordinarily extreme rainfall
384 event (or *outlier*). The inclusion of an *outlier*, with a recur-
385 rence period much longer than the length of the series, could
386 cause an anomalous effect in the calculated mean and standard
387 deviation values (Hershfield 1961). In a previous work (Casas
388 et al. 2008), the 1-day PMP values at 145 locations in Catalo-
389 nia were estimated following Hershfield's procedure. From
390 their annual maximum daily rainfall series, different statistical
391 parameters including frequency factors, k_m , were calculated.
392 Based on the 145 stations, the k_m values were plotted against
393 the mean \bar{x}_n , in order to consider an appropriate enveloping
394 curve that would give reliable estimates of 1-day PMP for
395 Catalonia (Rakhecha et al. 1992). One of the 145 stations used
396 was the Ebre Observatory, with a 46-year annual maximum
397 daily rainfall series obtained from the records of a totalizer rain
Q7 398 gauge between 1939 and 1985. The equation corresponds to
399 the enveloping curve ($k_m = -7.56 \ln \bar{x}_n + 40.5$) assigned a
400 frequency factor of $k_m=7.3$ to the mean daily rainfall value of
401 81.0 mm in the Ebre Observatory, resulting in an estimated 1-
402 day PMP of 396 mm for this station.

403 From the 98-year annual series of true maximum rainfall
404 amounts for durations between 5 min and 24 h recorded by the
405 siphon pluviograph, a new statistical estimation of the PMP in
406 the Ebre Observatory has been done. These are true maxima
407 because amounts were calculated using unrestricted time inter-
408 vals by any particular observation time, instead of the usual
409 fixed intervals between a beginning and an ending time.
410 Table 4 shows the statistical parameters and the observed fre-
411 quency factors for every considered duration. The mean value
412 of the annual maximum rainfall in 24 h registered by the
413 siphon gauge in the period 1905–2003 is 76.4 mm (see
414 Table 4). The WMO (1986) recommends a multiplicative

factor of 1.13 for annual maximum rainfall amounts registered
in a single fixed time interval to yield values closely approx-
imating those obtained from analysis based on true maxima.
Thus, considering this adjusting factor, the mean true rainfall
amount of 76.4 mm in 24 h would approximately correspond
to an amount of 67.6 mm for a 24-h fixed interval. The
enveloping curve obtained by Casas et al. (2008) assigns to
this mean value a frequency factor of 8.6, higher than any of
the observed frequency factors varying from 3.3 and 6.3
(Table 4). The highest frequency factors have been found for
durations between 3 and 15 h, presenting a maximum of 6.3
for approximately 9 h. This result seems to be related to the
meteorological scale of the most severe rain events registered
in our region, which are generally related to some of the fol-
lowing mechanisms: deep convection, very active fronts, oro-
graphic uplift, mesoscale convective systems, and, very often,
the presence of a low-level cyclonic center (Jansà et al. 2001).
All these meteorological organizations, producing precipita-
tion with very high efficiency, are included typically into the
 β -mesoscale with characteristic durations from 2 and 10 h. As
the purpose of our work is to obtain PMP, the maximum value
8.6 has been considered as a maximized frequency factor for
any duration. In fact, for the same location, the frequency
factor k_m for an established return period should be expected
not to depend on the duration (Koutsoyiannis 1999). Using
Eq. (8), the PMP for every duration t has been calculated
(Table 5, Fig. 9). The uncertainty of the PMP values has been
determined using a method proposed by Salas et al. (2014) for
statistical estimations of the PMP based on Hershfield's tech-
nique. This uncertainty has resulted lower than 10 % for each

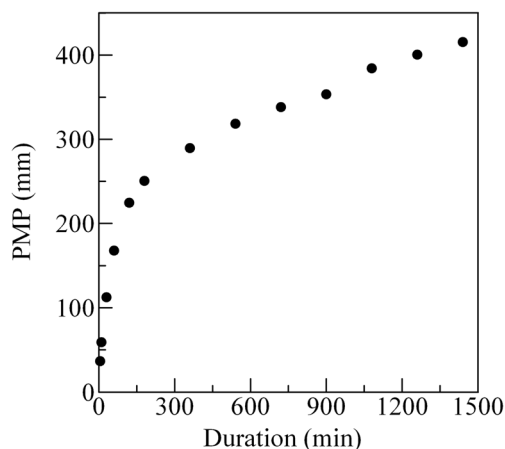


Fig. 9 PMP values for durations between 5 min and 24 h

445 of the considered durations. The value of the 1-day PMP has
 446 been estimated as 416 ± 38 mm, a range in which the value of
 447 396 mm found by Casas et al. (2008) for the same location is
 448 included.

449 **7 Conclusions**

450 The maximum rainfall amounts corresponding to time inter-
 451 vals between 5 min and 24 h, obtained from the digitalization
 452 of the strip charts recorded by two siphon rain gauges located
 453 in the Ebre Observatory between 1905 and 2003, have been
 454 satisfactorily fitted to a GPA distribution function. The master
 455 equation of the IDF curves in the location $I(t, T) =$

$$\frac{11.9 \log T + 9.9}{(6.9 T^{0.19} + t)^{0.75}}$$
 has been obtained.

456 Extreme rainfall events recorded in the Ebre Observatory
 457 showing a return period equal to or longer than 10 years for
 458 any of the considered durations from 5 min to 24 h have been
 459 selected. Seventy-one percent of these events occurred in au-
 460 tumn. Using a cluster analysis technique, these events have
 461 been characterized and classified into four clearly differenti-
 462 ed groups. The first group contains high-intensity storms with
 463 very short durations (equal to or shorter than 15 min),
 464 representing the local rainfall events showing a clear seasonal
 465 influence and diurnal cycle. The second group corresponds to
 466 the typical mesoscale durations, whereas synoptic rainfall
 467 events with intensities exceeding the 10-year return period
 468 level only for time intervals higher than 9 h constitute the third
 469 group. Finally, a rainfall pattern showing high rates for a large
 470 range of durations from 10 min to 24 h has been found, asso-
 471 ciated to local, meso-, and synoptic scale meteorological pro-
 472 cesses acting together. The synoptic situations corresponding
 473 to this kind of complex events (14 % of the studied cases) are
 474 characterized by the presence of a depression located SW of
 475 the Iberian Peninsula and a powerful blocking anticyclone on
 476 the European continent. In these situations, the prevailing
 477 wind on Catalonia at 850 hPa and on the surface is from the
 478 SE, resulting in a moist and warm advection from the
 479 Mediterranean Sea. These situations produce long-lasting ad-
 480 vective precipitation at the same time that convective phenom-
 481 ena linked to the SE warm flow in low levels cause extreme
 482 intense rainfall. A weighted intensity index taking into account
 483 the maximum rainfall rate in the 5-min, 1-h, 2-h, and 24-h
 484 time intervals has been calculated for every extreme rainfall
 485 event in order to reflect the contribution of every meteorolog-
 486 ical scale in the origin of rainfall and their complexity.

488 The PMP in the location for durations between 5 min and
 489 24 h has been estimated using a statistical technique. A value
 490 of 415 (± 38 mm) has been estimated for the 1-day PMP, very
 491 similar to the value of 396 mm found by Casas et al. (2008) for
 492 the same location.

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- Q6. SLP was defined as "sea level pressure." Please check if the definition is correct.
- Q7. Please check if the modifications are appropriate in the sentence "The equation corresponds to the enveloping curve...."
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