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Fe-Ti-Zr metasomatism in the oceanic mantle due to extreme differentiation of tholeiitic melts (Moa-Baracoa ophiolite, Cuba)

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ABSTRACT

Ti-rich amphibole, Mg-rich ilmenite, baddeleyite, zirconolite, srilankite, and zircon are important high-field-strength elements (HFSE) bearing phases in the Potosí chromitite bodies located in the Moho Transition Zone of the Cretaceous Moa-Baracoa suprasubduction zone ophiolite (eastern Cuba). Such HFSE-bearing phases were found in the interaction zone between gabbroic intrusions and chromitite pods. In addition to HFSE-bearing minerals, the studied samples are composed of Fe^{3+} and Ti-rich chromite, olivine (Fo₈₆₋₉₀), clinopyroxene (En₄₄₋₄₉), plagioclase (An₅₁₋₅₆), orthopyroxene (En₈₄₋₉₄), F-rich apatite, and Fe-Cu-Ni sulfides. The studied ilmenite hosting Zr oxides (baddeleyite, zirconolite, and srilankite) contains up to 13 wt.% MgO. The Potosí zirconolite is the first record of this mineral in ophiolitic chromitites and non-metamorphic ophiolite units, and it has relatively high REE contents (up to 10 wt.% of REE₂O₃) and the highest concentrations in Y₂O₃ (up to 11 wt.%) reported so far in zirconolite from terrestrial occurrences. Zircon is observed forming coronas surrounding ilmenite grains in contact with silicate minerals, and is characterized by very low U and Pb contents. The zircons formed after high temperature Zr diffusion in ilmenite (exsolution) and a subsequent reaction along grain boundaries following crystallization. Finally, U-Pb dating of baddeleyite exsolutions within ilmenite yielded an average age of 134.4 ± 14 Ma, which provides the first ever dating for a metasomatic event in Potosi that matches well (within uncertainty) the formation age of the oceanic crust of the eastern Cuba ophiolite. We propose that the occurrence of HFSE- and REE-bearing minerals in the Potosí chromite deposit is the result of a two stage process: first, water-rich and HFSE-rich residual melts are produced by intercumulus crystal fractionation after an evolved MORB (BABB)-like melt; and secondly, these residual melt fractions escaped the solidifying mush and extensively reacted and metasomatized the surrounding chromitites, crystallizing HFSE- and REE-bearing minerals and Fe-Cu-Ni sulfides.

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1. Introduction

Mantle metasomatism is a process by which the infiltration of melts or fluids through the Earth's upper mantle produces chemical and/or mineralogical modification of peridotites and associated rocks (Bodinier et al., 1990; Constantin, 1999; Menzies and Hawkesworth, 1987). This process leads either to the removal or addition of components of the original rock to varying extents, depending on the melt-fluid/rock ratio, the compositional nature of the metasomatic agent and the mode of infiltration (i.e., porous flow versus open fractures; O'Reilly and Griffin, 2013). As a result of variable melt/rock ratio

fingerprints in the infiltrated rocks such as newly formed minerals (Bedini et al., 1997; Bodinier et al., 1990). These minerals may be distinct to those originally forming the peridotite, giving rise to the so-called modal metasomatism, or alternatively they are identical to those already existing in the upper mantle rock, which produce the so-called stealth metasomatism (O'Reilly and Griffin, 2013). Recent advances of in-situ microanalytical techniques such as laser ablation ICPMS have allowed to better define the third style of metasomatism, known as cryptic metasomatism, produced under low melt/rock ratios and characterized by the modification of the distribution of major, minor, and trace elements in peridotite-forming minerals (O'Reilly and Griffin, 2013 and references therein). To date, most efforts in understanding mantle metasomatism have been focused on the study of ultramafic xenoliths and exhumed ultramafic massifs that have sampled the subcontinental lithospheric mantle (SCLM) underlying cratonic and off-craton regions of the Earth (Bodinier et al., 2008; Coltorti et al., 2004; Le Roux et al., 2007; Lenoir et al., 2001;

tios through the upper mantle, metasomatism leaves behind different

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Soustelle et al., 2009). In the SCLM, these three styles of metasomatism are very well preserved and have been identified to originate via a variety of metasomatic agents, including supercritical fluids and a variety of silicate liquids including silica-saturated, high-Na and high-K alkaline basaltic, as well as kimberlitic and carbonatitic compositions (e.g., Bodinier et al., 1990; Ionov et al., 1999; Coltorti et al., 1999, 2000; Greégoire et al., 2000; Laurora et al., 2001; Dawson, 2002; Lorand et al., 2004; Scamberulli et al., 2009; González-Jiménez et al., 2014a; Akizawa et al., 2017; Tassara et al., 2018). However, modal metasomatism is relatively underexplored in the mantle rocks of "normal oceanic lithosphere", as numerous studies focused on oceanic islands such as Hawaii, Kerguelen or Canary islands.

Reactive melt percolation, partial dissolution, and melt crystallization are common processes in extensional settings, causing large compositional variations in the lithospheric mantle (Bodinier and Godard, 2003; Kelemen et al., 1995a, 1995b; Rampone et al., 2018; Warren and Shimizu, 2010 and references therein). The resulting lithologies in the thermal boundary layer (TBL) and in the lithospheric mantle include replacive harzburgites, dunites, and impregnated peridotites with plagioclase and pyroxene (Basch et al., 2019a, 2019b; Rampone et al., 2018 and references therein). Ascending melts also lead to the formation of gabbroic rocks (olivine-gabbro, gabbro, gabbronorite, and ferrogabbro or oxide gabbro), often emplaced as intrusive dikes crosscutting peridotites in slow- (e.g., Morishita et al., 2004) and fast-spreading environments (e.g., Constantin, 1999). Morishita et al. (2004) reported an example of extreme metasomatism in the slow-spreading SW Indian Ridge, where extremely differentiated mantle-derived melts infiltrate oceanic peridotites thus forming Fe-Ti-rich gabbroic veins with a complex mineral assemblage including srilankite, Mg-rich ilmenite, apatite, and rutile.

Similar processes, such as multi-stage melt infiltration and magma-peridotite interaction, are recorded in many ophiolites (Kelemen et al., 1992, 1995b; Python and Ceuleneer, 2003), especially in the Moho Transition Zone (hereafter MTZ) of ophiolitic sequences. The MTZ, located between residual peridotites and layered crustal gabbros, is characterized by variable thickness but significant lateral continuity of a rock assemblage made of harzburgites, dunites, chromitites, wehrlites, troctolites, pyroxenites, in addition to gabbroic sills and dikes. Many bodies of dunite, wehrlite, troctolite, and websterite originated from impregnation of residual peridotites (e.g., Marchesi et al., 2006), while extensive melt-rock interaction by reactive porous flow leads to selective enrichment in incompatible elements (Basch et al., 2019a).

Eastern Cuban ophiolites are characterized by large exposures of the MTZ, where interaction between residual peridotites and percolating melts have resulted in the formation of dunite bodies and associated chromitites (Marchesi et al., 2006; Proenza et al., 1999a). In the Potosi chromite deposit, Proenza et al. (2001) reported a very peculiar mineral assemblage made up of Fe-Ni-Cu sulfides, apatite, amphibole, and unidentified oxides of Ti, Zr, Y, Ca, Fe, Hf, and REE along the contact between the chromitite pods and intruding gabbroic dikes.

The present study explores in detail the metasomatic assemblage and major and trace element composition of uncommon HFSE-bearing minerals (Mg-rich ilmenite, Hf-bearing baddeleyite, REE-rich zirconolite, srilankite, zircon) found in the interaction zone between gabbroic intrusions and chromitite pods. This is an unprecedented scenario where a highly evolved tholeiitic melt interacted with chromitite in the MTZ, producing a unique metasomatic assemblage enriched in Fe, Ti, Zr, and other high-field-strength elements (HFSE).

2. Geological setting

The Mesozoic Cuban ophiolite belt, that outcrops discontinuously over more than 1,000 km along length of the island (Fig. 1a), represents the largest exposure of ophiolites in the Caribbean region (Garcia-Casco et al., 2006; Iturralde-Vinent, 1996; Iturralde-Vinent et al., 2016; Lewis et al., 2006). This belt represents slices of oceanic lithosphere obducted during Latest Cretaceous to Eocene times as a consequence of the collision of the Caribbean volcanic arc with the Jurassic-Cretaceous passive margins of the continental Maya block and the Bahamas platform (Iturralde-Vinent et al., 2006, 2016; Garcia-Casco et al., 2008a).

In Eastern Cuba, ophiolitic bodies are grouped into the Mayarí-Baracoa Ophiolite Belt (MBOB). This region is characterized by NE-directed tectonic nappes accreted in the Late Cretaceous - Earliest Paleocene (Cobiella-Reguera, 2005; Iturralde-Vinent et al., 2006). These nappes include Early to Late Cretaceous oceanic arc sequences (Marchesi et al., 2007; Proenza et al., 2006), some of them subducted during the latest Cretaceous (Garcia-Casco et al., 2008a and references therein), the MBOB (Proenza et al., 1999b) and closely associated Early Cretaceous low-P and high-P (subduction-related) serpentinite-matrix mélanges (Garcia-Casco et al., 2008b; Blanco-Quintero et al., 2010, 2011; Lázaro et al., 2009, 2016; Cárdenas-Párraga et al., 2017), a Late Cretaceous metamorphic sole (Lázaro et al., 2013, 2015), and high-pressure metamorphosed passive-margin sediments of the Caribbeana terrane (Garcia-Casco et al., 2008a and references therein). The MBOB (170 km long, 10 - 30 km wide) comprises two allochthonous massifs separated by major fault zones: the Mayarí-Cristal (exposing 5 km of mantle section) to the west and the Moa-Baracoa (2.2 km of peridotites) to the east (Fig. 1b; Proenza et al., 1999a, 1999b; Marchesi et al., 2006). The ophiolite includes the mantle and crustal sequences, is highly dismembered, and represents MOR-like lithosphere formed in the suprasubduction zone setting of the intra-oceanic Caribbean arc (Gervilla et al., 2005; Marchesi et al., 2006, 2007; Proenza et al., 1999b, 2018). More than 250 chromite deposits and occurrences have been described in the MBOB (Gervilla et al., 2005; González-Jiménez et al., 2011; Proenza et al., 1999b). They are divided into three mining districts according to the composition of chromite forming the chromitite ore: Mayarí (high-Cr), Sagua de Tánamo (high-Cr and high-Al) and Moa-Baracoa (high-Al) (Fig. 1b).

2.1. The Moa-Baracoa ophiolitic massif

The Moa-Baracoa massif consists of mantle tectonite harzburgite with subordinate dunite and a MTZ associated with layered gabbros (about 500 m thick) and mafic volcanic rocks (Fig. 1c; Marchesi et al., 2006). Towards the top of the mantle tectonites, harzburgite contains increasing amounts of dunite, gabbro sills, and chromitite, all forming elongated pseudotabular bodies oriented parallel to the foliation of the host harzburgite, as well as discordant dikes of wehrlite, troctolite, olivine gabbro, and pegmatitic gabbro (Flint et al., 1948; Guild, 1947; Marchesi et al., 2006; Proenza et al., 1999a, 1999b, 2001). The peridotites and layered gabbros are in tectonic contact with the pillow basalts of the Morel formation (88 - 91 Ma; Iturralde-Vinent et al., 2006), which show a back-arc geochemical affinity. Marchesi et al. (2006) established the genetic link between the Morel formation and the cumulate gabbros of the Moa-Baracoa massif. Therefore, the Moa-Baracoa Ophiolite Massif is interpreted to have formed in a back-arc environment (Gervilla et al., 2005; Lázaro et al., 2015; Proenza et al., 2006).



Fig. 1. (a) Generalized geologic-tectonic map of Cuba modified from Iturralde-Vinent et al. (2016); (b) Geological map of the Mayarí-Baracoa Ophiolitic Belt (MBOB) (modified from Pushcharovsky et al., 1988) showing the location of the Potosí deposit; (c) Schematic lithostratigraphic column of Moa-Baracoa with location of Potosí (modified from Marchesi et al., 2006).

2.2. The Potosí chromitites and the gabbroic intrusions

The Potosí chromitites are located 800 m deeper than the layered gabbros forming the MTZ at the Moa-Baracoa Ophiolitic Massif (Fig. 1c) (Proenza et al., 2001). They form lenses of variable sizes and irregular shapes (tabular to lenticular) hosted in dunite within harzburgite. The bodies are oriented N 40° E to N 74° E with dips between 32° and 45° to the NW, usually filling the foliation of the host peridotites. Several generations of gabbroic intrusions crosscut the chromitite bodies, usually following pull-apart fractures with brecciated zones at their margins (Fig. 2a-d; Proenza et al., 2001). Dikes have highly variable thicknesses, from a few centimeters to 2 m across and locally show a distinct grain-size zoning with pegmatitic textures in the dike cores that reach grain sizes larger than 15 cm (Fig. 2c-e) and finer grain sizes closer to the contacts with chromitite. Disrupted fragments of chromitite occur as inclusions within the gabbroic intrusions (Fig. 2d). The contacts between the chromitites and the gabbroic intrusions may be either neat (Fig. 2e) or gradational, the latter characterized by a brecciated zone where chromite is included within the gabbroic dikes (Fig. 2f)

Chromitite in Potosí can be classified as: i) ordinary podiform chromitite, preexistent to gabbroic intrusions, or ii) metasomatized chromitite, associated with or intruded by gabbroic dikes and denoted as "brecciated chromite ore" and "sulfide-rich chromite ore" by Proenza et al. (2001). Ordinary chromitites represent more than 90 vol.% of the ore body and show mineralogical, textural and chemical features similar to other Al-rich chromitite bodies in the MBOB (Gervilla et al., 2005; Proenza et al., 1999b) and ophiolitic complexes worldwide (González-Jiménez et al., 2014b; Arai and Miura, 2016 and references in these papers). Metasomatized chromitite is composed mostly of recrystallized Ti- and Fe^{3+} - rich chromite, Ti-rich amphibole, orthopyroxene, clinopyroxene, plagioclase, and Fe-Cu-Ni sulfides, the latter formed by fractionation of immiscible sulfide liquid segregated from the mafic intrusions (Proenza et al., 2001).

3. Studied samples and analytical techniques

3.1. Studied samples

A total of 20 hand samples and 38 thin sections were examined for this study (Appendix 1), including ordinary chromitite and its host peridotite (5 samples), metasomatized chromitite (11 samples), and gabbroic dikes (4 samples) (Fig. 2a-b). Most samples were taken from the outcrops of old mine works and some were collected in the mine dumps.

3.2. Analytical techniques

Modal composition of the POT-2 sample was calculated using point counting on thin section with the software *JMicroVision*. Whole rock geochemical analyses were performed at the Centre for Scientific Instrumentation (CIC) of the University of Granada. Samples were carefully prepared by removing secondary veins and weather-

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Fig. 2. (a) Schematic cross-section of a chromitite body in Potosi crosscut by multiple generations of gabbroic dikes. Red dots indicate the location of the samples; (b) Detail of the interaction zone between the chromitite and the intruding gabbroic dike. Brighter blue indicates differentiated melts. Red dots indicate the location of the samples; (c) Field relations of a gabbroic dike intruding the peridotite. Note the variation in grain size and the area with pegmatitic texture; (d) Gabbroic dike, with pegmatitic textures, intruding chromitite. Note the disrupted fragments of chromitite as inclusions within the gabbroic dike; (e) Hand sample of the contact between the gabbroic dike and the chromitite; and (f) Hand sample of brecciated chromitite. Abbreviations: Chr – chromite, Pl – plagioclase, Cpx – clinopyroxene.

ing products before crushing and powdering in an agate mill. Major element and Zr concentrations were determined by X-ray fluorescence (XRF) spectrometry on glass beads (~ 0.6 g of powdered sample diluted in 6 g of $Li_2B_4O_7$) using a Philips Magix Pro (PW-2240). Precision was better than $\pm 1.5\%$ for an analyte concentration of 10 wt.% and the precision for Zr was better than $\pm 4\%$ at 100 ppm concentration. The analyses were recalculated to an anhydrous basis (100 wt.%). Trace elements were analyzed by ICP Mass Spectroscopy (ICP-MS) at the same institution. ~ 100 mg of sample were digested with HNO₃ + HF in a Teflon lined vessel at 180 °C and ~ 200 psi for 30 min, evaporated to dryness, and subsequently dissolved in 100 ml of 4 vol.% HNO₃. During analytical sessions, the procedural blanks and reference materials PMS, WSE, UBN, BEN, BR, and AGV were run as unknowns and primary standards to obtain concentrations (Govindaraju, 1994). Precision was better than $\pm 2\%$ and $\pm 5\%$ for analyte concentrations of 50 and 5 ppm, respectively. Major and trace element bulk analyses results are presented in Appendix 2.

Polished thin sections were studied in detail by optical microscopy and scanning electron microscopy, using both a Quanta 200 FEI XTE 325/D8395 scanning electron microscope (SEM) and a JEOL JSM-7100 field-emission SEM at the Universitat de Barcelona (CCiTUB). Operating conditions were 15 - 20 kV accelerating voltage and 5 nA beam current. Quantitative electron microprobe analyses (EMPA) were also conducted at the CCiTUB using both CAMECA SX50 and JEOL JXA-8230 electron microprobes, operating in wavelength-dispersive spectroscopy (WDS) mode. The analytical conditions are described in Appendix 3.

Trace element compositions of chromite, ilmenite, clinopyroxene, orthopyroxene, olivine, amphibole, apatite, zircon, and baddeleyite were measured in-situ on polished thin sections ($30 \mu m$) by an induc-

tively coupled mass spectrometer Agilent 8800 QQQ ICP-MS interfaced to a laser ablation extraction line Photon Machines Analyte Excite 193 at the Instituto Andaluz de Ciencias de la Tierra (CSIC-UGR, Granada). The analytical conditions are described in Appendix 3. Six grains of baddeleyite were analyzed for U/Pb isotopes with the same equipment at the same institution. A spot diameter of 30 μ m was used with the same analytical conditions as for zircon. Phalaborwa baddeleyite was used as a reference material and in-house Mogok baddeleyite for secondary/validation. Data (Appendix 4) were reduced using Igor Pro iolite 3.6 software, VisualAge DR, ET-Redux, and Isoplot for data presentation.

4. Results

4.1. Petrological features of the gabbroic intrusions

Two types of gabbroic intrusions have been characterized in Potosi (Fig. 3). The first type (POT-2) corresponds to a melagabbro (Fig. 3a) containing olivine (64%), clinopyroxene (21%), plagioclase (14%), orthopyroxene (< 1%), and accessory oxides with a general coarse grain size. Olivine and plagioclase are cumulus minerals and clinopyroxene is intercumulus (mesocumulate texture). Clinopyroxene reaches up to 2 cm in length and develops a poikilitic texture. The major and trace element composition of this type of intrusion (Appendix 2) is similar to the primitive sills and dikes crosscutting the MTZ in the Moa-Baracoa massif reported by Marchesi et al. (2006) (Fig. 3b, c). The trace element contents are slightly richer than chondrite and the chondrite-normalized REE pattern shows a slightly positive slope from LREE to MREE and a flat pattern from MREE to HREE (Fig. 3d, e).



Fig. 3. (a) Modal compositions of the Potosi studied gabbroic dikes compared with gabbro sills and dikes, and layered gabbros from the Moa Baracoa Massif (data from Marchesi et al., 2006); (b-c) Whole rock major and trace element compositions of the Potosi studied gabbroic dikes compared with gabbro sills and dikes, and layered gabbros from the Moa Baracoa Massif (data from Marchesi et al., 2006); (b) Al_2O_3 (wt.%) versus MgO (wt.%), and (c) Zr (ppm) versus TiO_2 (wt.%). All data on anhydrous basis in wt.%; (d) Chondrite-normalized REE patterns of the studied Potosi gabbroic dikes compared with gabbro sills and dikes, and layered gabbros from the Moa Baracoa Massif (data from Marchesi et al., 2006). Normalizing values from Sun and McDonough (1989); and (e) Primitive mantle-normalized patterns of the studied Potosi gabbroic sills and dikes, and layered gabbros from the Moa Baracoa Massif (data from Marchesi et al., 2006). Normalizing values from Sun and McDonough (1989).

The second type of gabbroic intrusion (POT-1) corresponds to a leucogabbro dike with pegmatitic texture (Fig. 2e, 3a) containing plagioclase (70%) and clinopyroxene (28%) crystals up to 3 cm in length with accessory oxides (around 2%). The major element composition of the dike (Appendix 2) is similar to the more evolved gabbroic sills and dikes (Fig. 3b) of the Moa-Baracoa massif (Marchesi et al., 2006). However, it shows an important enrichment in trace elements (Appendix 2), notably in Zr (Fig. 3c). The chondrite-normalized REE (Fig. 3d) and primitive mantle-normalized trace element (Fig. 3e) patterns show enrichment by one order of magnitude compared to those of the Moa-Baracoa gabbroic dikes, sills, and layered gabbros (Marchesi et al., 2006).

Plagioclase crystals (up to 2 cm) in the gabbro dike show polysynthetic twinning (Fig. 4a) and partial alteration to (clino)zoisite and fine-grained phyllosilicate aggregates. Fresh plagioclase shows labradorite composition ($An_{51} - An_{56}$) (Appendix 5) with lim-



Fig. 4. (a) Transmitted light (crossed nicols) photomicrograph of chromite with amphibole rim and partially altered plagioclase; (b) Transmitted light (crossed nicols) photomicrograph of clinopyroxene 1 (core) and clinopyroxene 2 (rim) surrounded by alteration minerals (probably chlorite); (c) Petrographic thin section of the interaction zone between chromitite and gabbroic dikes, where the main mineralogy includes recrystallized (metasomatized) chromite, amphibole and plagioclase; (d) Petrographic thin section of the interaction zone between chromitite and gabbroic dike, where the main mineralogy includes recrystallized (metasomatized) chromite, plagioclase and Fe-Ni sulfides; (e) Transmitted light (crossed nicols) photomicrograph of recrystallized chromite and ilmenite (opaque) surrounded by an amphibole rim with host plagioclase and alteration minerals; (f) Transmitted light (particular) photomicrograph of recrystallized chromite and ilmenite (opaque) surrounded by an amphibole hosting multiple oxide inclusions along the cleavage planes; (g) Reflected light photomicrograph of (f)—the opaque minerals in (f) are recrystallized chromite and ilmenite; (h) Reflected light photomicrograph of chromite hosting multiple oriented ight (s) photomicrograph of chromite hosting multiple oriented oriented and ilmenite (and petitod); (i) Transmitted light (crossed nicols) photomicrograph of orthopyroxene in contact with sulfides (pyrrhotite and petitod); (i) Transmitted light (crossed nicols) photomicrograph of orthopyroxene in contact with fine-grained olivine within recrystallized chromite. Abbreviations: Chr – chromite, Pl – plagioclase, Amp – amphibole, Cpx – clinopyroxene, Chl – chlorite, Ilm – ilmenite, Ol – olivine, Pn – pentlandite, Po – pyrrhotite, Opx – orthopyroxene.

ited compositional variation at the grain scale. Clinopyroxene crystals (up to 3 cm) show deformation revealed by folding of the cleavage planes. These crystals host many elongated oxide inclusions along the cleavage planes and show exsolution of orthopyroxene lamellae. The composition of clinopyroxene is relatively narrow, corresponding to the limit between diopside and augite $(En_{44.49}Wo_{48.46}Fs_{8.5})$, with Mg#=0.85 – 0.92 and high TiO₂ (0.36 – 1.05 wt.%) (Appendix 5). Some clinopyroxene grains show rims with slightly different extinction angles (Fig. 4b) and higher diopside contents.

4.2. Metasomatized chromitite

The interaction between the pegmatitic gabbro enriched in trace elements and the ordinary chromitite generated unusual mineral phases in the metasomatized chromitite. This type of rock shows strong variability in textures and modal mineralogy at the hand-sample and thin-section scales. Most samples show coarse granular textures with crystals up to 2 cm in length (Fig. 2e, f, 4c, d). The primary igneous-metasomatic phases are Fe^{3+} -Ti-rich chromite, Ti-rich amphibole, orthopyroxene, olivine, Mg-rich ilmenite, and Fe-Cu-Ni sulfides (Fig 4c, d). This assemblage experienced variable late overprinting that is attributed to ocean floor alteration/metamorphism, which produced the secondary serpentine-group minerals, chlorite, ferrian-chromite, magnetite, actinolite-tremolite, rutile, and titanite.

Metasomatized chromite forms generally subhedral crystals with sizes ranging between 0.5 and 2 cm (Fig. 4c-e). Locally, octahedral chromite crystals are cemented by sulfides. Chromite is usually associated with amphibole and ilmenite (Fig. 4f, g) and most chromite grains host abundant crystallographically-oriented ilmenite lamellae (Fig. 4h). Metasomatized chromite has higher Cr# [Cr/(Cr+Al)] (0.57 – 0.70), higher Fe^{3+} [Fe^{3+} /(Fe^{3+} +Cr+Al)] (0.10 – 0.30) and lower Mg# [Mg/(Mg+Fe²⁺)] (0.31 – 0.53) than chromite from ordi-

nary chromitites (Cr#=0.44 – 0.54, Fe³⁺#< 0.05, and Mg#=0.66 – 0.74; see also Proenza et al., 2001). It is also enriched in TiO₂ (1.02 – 4.34 wt.%), Fe³⁺ (up to calculated 21.5 wt.% Fe₂O₃), Fe²⁺ (up to calculated 25.31 wt.% FeO), and V₂O₃ (up to 0.7 wt.%) (Fig. 5) (Appendix 6). Two types of metasomatized chromite can be distinguished (Fig. 5):

- (I) Type I bears an intermediate composition between ordinary and metasomatized s.s. chromite (type II, see below) (Cr#=0.57 0.58, Fe³⁺#=0.10 0.13, and Mg#=0.51 0.53), which roughly corresponds to the "brecciated chromite ore" defined by Proenza et al. (2001).
- (II) Type II (Cr#=0.62-0.70, Fe³⁺#=0.19-0.30, and Mg#=0.31-0.39) further interacted with the gabbroic dikes and corresponds to the "sulfide-rich chromite ore" (10-30 vol.% Fe-Ni-Cu sulfides) of Proenza et al. (2001).

The change in composition from ordinary to type II metasomatized s.s. chromite is defined by an enrichment trend in Cr#, $Fe^{3+}#$, and TiO_2 (Fig. 5). In type II metasomatized chromitite the sulfides are usually intergranular to (or forming rims around) chromite (Fig. 4d, h) and are mainly pyrrhotite, pentlandite, chalcopyrite and cubanite, with minor amounts of chalcocite and valleriite (Proenza et al., 2001).



Fig. 5. Composition of chromite from Potosí, (a) Variation of Cr# [Cr/(Cr+Al)] versus Mg# [Mg/(Mg+Fe²⁺)]. The areas represent the compositions of the Potosí ordinary (sulfide-poor chromite ore) and metasomatized chromite (brecciated and sulfide-rich chromite ores) reported by Proenza et al. (2001), and the dashed area is the compositional field of the chromitites from the Moa-Baracoa Massif (Proenza et al., 1999b); (b) TiO₂ versus Fe³⁺# [Fe³⁺/(Fe³⁺+Cr+Al)]; (c) MnO versus Fe³⁺#; and (d) V₂O₃ versus Fe³⁺#.

Brown amphibole forms large crystals (up to 2 cm) and coronas around the chromite (Fig. 4a, e). Large crystals usually host a myriad of elongated inclusions (up to 15 μ m long) of ilmenite, rutile and titanite following the cleavage planes (Fig. 4f, g). Amphibole composition (Fig. 6a-h) is Ti-rich with TiO₂ ranging between 2.41 and 5.46 wt.%, Mg# between 0.82 and 0.94, Cl up to 0.05 wt.% and F up to 0.5 wt.%, corresponding mainly to Ti-rich magnesio-hastingsite and, to a lesser extent, Ti-rich pargasite and ferri-kaersutite compositions (calculations following Locock, 2014 and the classification scheme of Hawthorne et al., 2012; Fig. 6a-b; Appendix 5). Orthopyroxene is not abundant and forms grains up to 3 mm that are generally free of inclusions and devoid of zoning with an enstatite composition ($En_{84.94}Wo_{4.1}Fs_{12.5}$), with Mg#=0.87 – 0.89 and high TiO₂ contents (0.23 – 0.47 wt.%, Appendix 5). These orthopyroxenes are in contact with rare fine-grained Fo_{86.91} olivine (up to 200 µm) (Fig. 4i) with NiO ranging between 0.12 and 0.22 wt.% (Appendix 5). The higher forsterite contents correspond to typical mantle olivine, including olivine from the regional Moa-Baracoa chromities (Fo_{90.97}; Proenza et al., 1999b), while the lower forsterite values are closer to those of gabbroic intrusions (Fo_{74.82}; Marchesi et al., 2006



Fig. 6. (a-b) Amphibole classification diagrams according to Hawthorne et al., 2012. (a) Variation of ${}^{c}[Mg/(Mg+Fe^{2^+})]$ a.p.f.u. versus ${}^{c}[Fe^{3^+}/(Fe^{3^+}+Al)]$ a.p.f.u.; (b) ${}^{c}[Fe^{3^+}/(Fe^{3^+}+Al)]$ a.p.f.u. versus ${}^{c}[Ti]$ a.p.f.u; (c-h) Major element variations in amphiboles from Potosi. The fields corresponding to "oceanic and continental lithospheric veins", "continental basanite xenocrysts", and "peridotites from xenoliths and ophiolites" are from Pilet et al. (2008). (c) Variation of SiO₂ (wt.%) versus Mg# [Mg/(Mg+Fe)]; (d) Variation of TiO₂ (wt.%) versus Mg#; (e) Variation of Al₂O₃ (wt.%) versus Mg#; (f) Variation of CaO (wt.%) versus Mg#; (g) Variation of Na₂O (wt.%) versus Mg#; and (h) Variation of K₂O (wt.%) versus Mg#.

). Olivine crystals locally include elongated grains of oriented Cr-spinel and ilmenite.

Ilmenite forms subhedral to anhedral crystals (up to 1 cm) located at the rims of (Fig. 4g), or within, chromite crystals (Fig. 7a, d, f, h) and forms oriented lamellae in chromite following its crystallographic orientation (Fig. 4h). The content in the geikielite component (Appendix 7) varies between 1 and 45 mol% MgTiO₃ (averaging 26 mol%) (Fig. 8a, b). Ilmenite grains are devoid of chemical zoning and compositional variations are independent of textural position, as illustrated by strikingly different geikielitic contents in adjacent grains of ilmenite (Fig. 8a, b).

Several mineral phases occur as inclusions within, or at the rims of, grains of ilmenite. The most abundant is apatite, forming euhedral to subhedral elongated crystals up to 500 μ m long (Fig. 9a, b) in contact with or included within ilmenite. Seldom, apatite crystals include sulfide inclusions (Fig. 9b). Apatite contains 0.85 – 1.52 wt.% F and 0.54 – 0.77 wt.% Cl, with a Cl/F (molar) ratio between 0.33 and 0.48 (Appendix 8). Relatively small (up to 500 μ m in length) ilmenite crystals in contact with chromite host accessory Zr-oxides (baddeleyite, zirconolite, and srilankite) (Fig. 7), whereas large ilmenite crystals (up to 0.5 cm long) in contact with silicates (Fig. 9) show zircon coronas. It is noteworthy that Zr-oxides and zircon do not coexist surrounding, or within, a given ilmenite crystal.

Zr oxides form subhedral to anhedral elongated crystals with sizes ranging between 10 and 80 µm long within ilmenite (Fig. 7). Occasionally, the Zr oxide inclusions have multiple mineral phases (e.g., baddelevite and zirconolite or baddelevite and srilankite; Fig. 7i). Baddelevite crystals show Hf-enrichment (HfO₂ between 1.52 and 1.96 wt.%) and low Nb, Nd, and U contents (up to 0.06 wt.% Nb₂O₃, up to 0.06 wt.% Nd₂O₃, up to 0.07 wt.% U₂O₃) (Appendix 8). Zirconolite composition deviates from the ideal CaZrTi₂O₇ (Fig. 10a), with very low U and Th contents (Fig. 10b), relative enrichment in REE $(\text{REE}_2\text{O}_3 = 9.25 - 10.7 \text{ wt.}\%)$ (Fig. 10c) and containing exceptionally high Y (up to 11.06 wt.% of Y₂O₃) (Appendix 9). The studied zirconolite contains the highest Y concentrations reported to date in terrestrial zirconolite and the observed dominant coupled substitution is: $2Ca^{2+}$ + $Ti^{4+} = 2(REE, Y)^{3+} + (Fe, Mg)^{2+}$ (Fig. 10d). Two compositional groups of zirconolite are observed with different degrees of coupled substitution (Fig. 10).

Zircon grains (up to 50 μ m) usually form rims around large ilmenite crystals, clustering at the contact with silicates (Fig. 9a, d), and they may appear elongated following the rim of ilmenite (Fig. 9e-g). Generally, zircon coronas show granular textures (Fig. 9f) and include ilmenite grains in between. Zircon is devoid of zoning (Appendices 10 and 11) with HfO₂ content ranging between 0.17 and 1.54 wt.%, which is within the range of abundance for natural zircon (0.5 – 5 wt.%; Belousova et al., 2002). These zircon coronas generally remain unaltered during alteration of ilmenite to rutile and titanite (Fig. 9h, i). Occasionally, zircon is replaced by an aggregate of fine-grained baddeleyite (secondary baddeleyite) together with Mg-Al silicates (Fig. 9c). Figure 11 summarizes the main petrographic features and mineral relations observed in the metasomatized chromitite in Potosí.

4.3. Mineral trace element compositions

Minerals from the metasomatized chromitite show relatively unusual trace element distribution and patterns (Appendix 12). Trace elements of chromite normalized to MORB chromite (Fig. 12a) show a positive anomaly in Ti and negative anomaly in Ni, which is not typical in Al-rich ophiolitic chromitites of the Moa-Baracoa massif, usually displaying relatively flat patterns (e.g., Colás et al., 2014). The chondrite-normalized REE patterns of Ti-rich amphibole are characterized by LREE depletion, nearly flat HREE and negative



Fig. 7. Backscattered electron images of: (a) Elongated ilmenite crystals within chromite. Note the bright inclusions within ilmenite that correspond to Zr-bearing oxides; (b) Zoom of (a) showing a baddeleyite inclusion within ilmenite in contact with chromite; (c) Second zoom of (a) showing a baddeleyite inclusion within ilmenite in contact with chromite; (d) Ilmenite subrounded crystal hosted in chromite; (e) Zoom of (d) showing elongated inclusions of baddeleyite and probably srilankite within ilmenite; (f) Ilmenite crystal at the border of chromite; (g) Zoom of (f) showing a zirconolite in clusion with minor srilankite in the upper part, hosted in ilmenite in the contact with chromite; (h) Ilmenite crystals in equilibrium with chromite hosting elongated baddeleyite together with zirconolite; and (j) Zoom of (h) showing an inclusion of baddeleyite inclusion within ilmenite. Abbreviations: Chr – chromite, Ilm – ilmenite, Bdy – baddeleyite, Srl – srilankite, Amp – amphibole, Ol – olivine, Opx – orthopyroxene, Zirc – zirconolite.



Fig. 8. (a) Ilmenite composition in terms of the pyrophanite (MnTiO₃)–ilmenite (FeTiO₃)–hematite (Fe₂O₃) endmembers. The compositional fields are from Tompkins and Haggerty (1985) and the shaded areas correspond to Morishita et al. (2004) in dark grey and Lorand and Gregoire (2010) in light grey; and (b) Ilmenite composition in terms of the hematite (Fe₂O₃)–ilmenite (FeTiO₃)–hematite (Fe₂O₃)–ilmenite (FeTiO₃)–hematite (Fe₂O₃) endmembers. Fields for kimberlitic ilmenites and basic rocks are from Mitchell (1977), the shaded areas correspond to Morishita et al. (2004) in dark grey and Lorand and Gregoire (2010) in light grey.

Eu anomalies (Fig. 12b), being similar in shape to (but more enriched than) associated clinopyroxene (Fig. 12c). Clinopyroxene shows two characteristic REE patterns corresponding to the cores and rims of grains. Clinopyroxene cores are particularly depleted in LREE and show strong fractionation from HREE to LREE (Fig. 12c), whereas clinopyroxene rims are characterized by higher trace element contents and larger negative Eu anomalies (Fig. 12c), with patterns depleted in LREE and a flat segment from MREE to HREE. Similar to clinopyroxene rims, orthopyroxene REE patterns (Fig. 12d) show strong REE fractionation, with a positive slope from LREE to HREE with negative Eu anomalies, but concentrations are one order of magnitude higher and the Eu negative anomaly is stronger.

Ilmenite trace element patterns (Fig. 12e) show positive anomalies in Nb and Ta and strong negative anomalies in Ni. Apatite shows REE patterns with a negative slope from LREE to HREE and slight Eu negative anomalies (Fig. 12f). The main substitution of Ca in the apatite structure is by Y and REE (REE+Y=161 – 405 ppm).

Zircon shows very low U (up to 54 ppm), Pb (up to 5 ppm), and Th contents (up to 6 ppm) (Appendix 12). Its crystallization temperature has been estimated using the Ti-in-zircon thermometer (Ferry and Watson, 2007; Watson et al., 2006), but due to the intimate intergrowth of zircon and ilmenite, the Ti content in nominally clean zircon reaches high values (> 4320 ppm) thus resulting in extremely high apparent crystallization temperatures that are not reasonable. The analyzed zircon grains have very pronounced Ce/Ce* (25 - 218) and the chondrite-normalized REE patterns (Fig. 12g) show a positive steep slope from LREE to HREE with positive Ce and negative Eu anomalies. The chondrite-normalized REE patterns of zirconolite show positive slopes from LREE to MREE and flat segments from MREE to HREE (Fig. 12h) with negative Eu anomalies.

4.4. U/Pb dating

Zircon grains forming coronas around ilmenite crystals are extremely depleted in U and Pb, especially in Pb^{207} , making it difficult to obtain a reliable date. The analyses yielded an average age (n=43) of 369 ± 24 Ma and a secondary cluster at 57 ± 17 Ma. On the other hand, analyses on 7 baddeleyite grains yielded one concordant age and 6 discordant but equivalent ages of 134.4 ± 14 Ma (Appendices 4 and 13).

5. Discussion

5.1. Meaning of the Fe-Ti-Zr minerals in the chromitite: contrasting literature data

We report metasomatic interaction of gabbroic melts with the chromitite located in the MTZ of the Moa-Baracoa ophiolite resulting in the metasomatism of pre-existing chromitite and the formation of a new exotic assemblage of Fe-Ti-Zr minerals. As noted above, the newly-formed mineral assemblage includes Ti-rich amphibole, Mg-rich ilmenite, Zr oxides, and zircon, and has characteristics that resemble those from SCLM settings, kimberlites, carbonatites or alkaline basalts (e.g., Lorand and Gregoire, 2010). Similar amphibole, usually K-rich, is typical of the MARID (mica-amphibole-rutile-ilmenite-diopside) assemblage in mantle xenoliths from kimberlites (e.g., Dawson and Smith, 1977; Fitzpayne et al., 2018; Grégoire et al., 2002), whereas Mg-rich ilmenite (Fig. 8) is also typical from kimberlites (e.g., Xu et al., 2018). Mg-ilmenite compositions have also been described in megacrysts in alkaline basalts from Algeria (Leblanc et al., 1982), in zirconolite-bearing ultrapotassic veins in mantle xenoliths from the Mt. Melbourne volcanic field in Antarctica (Hornig and Wörner, 1991), in some metasomatized peridotite xenoliths and peridotite massifs (Bodinier et al., 1996; Ionov et al., 1999), as well as in the amphibole-rich veins in the Lherz Massif (Lorand and Gregoire, 2010). The studied ilmenite in Potosí contains particularly low trivalent iron and is devoid of hematite exsolution, with compositions very similar to the ilmenite found in the amphibole veins of the Lherz Massif (Lorand and Gregoire, 2010) both in major (Fig. 8) and trace elements (Fig. 12e). In addition, the chromite trace elements (Fig. 12a) show similarities with accessory chromite from kimberlites and picrites (Yao, 1999).

Zirconolite is a frequent mineral in carbonatites, kimberlites, ultrapotassic metasomatic veins, skarns, as well as many others rock-forming settings (e.g., Williams and Gieré, 1996; Fig. 10). However, the only known occurrence of zirconolite in chromitite is in the Finero SCLM peridotites (Zaccarini et al., 2004). Most zirconolite re-



Fig. 9. Backscattered electron images of: (a) Zircon corona around ilmenite in contact with altered plagioclase. Ilmenite hosts a large apatite crystal; (b) Zoom of (a) showing a detail of the apatite crystal including Cu sulfides, and the zircon corona; (c) Zoom of (a) showing alteration of granular zircon to secondary baddeleyite mixed with silicates; (d) Large ilmenite crystal with zircon coronas along the ilmenite borders in contact with altered silicates. Note that in the region where ilmenite is in contact with sulfides, zircons are not present; (e) Zoom of (d) showing details of zircon grains surrounding ilmenite. Note the elongated shape of zircons; (f) Zoom of (d) showing the granular texture of zircons forming the coronas. Note the presence of interstitial ilmenite grains within these of zircons and that zircon grains are in contact with altered plagioclase; (g) Zoom of (d) showing detail of the zircon rims; (h) Ilmenite altered to rutile and titanite. The zircon corona around the primary ilmenite crystal remains immobile and unaltered during ilmenite replacement; and (i) Zircon corona around ilmenite defines ilmenite grain boundaries previous to alteration. Abbreviations: Ap - apatite, Pl - plagioclase, Ilm ilmenite, Zrn - zircon, Bdy2 - secondary baddeleyite, Chr - chromite, Rt - rutile, Ttn titanite

ported in the literature cluster in a similar compositional field (Fig. 10), but the Potosí zirconolite plots close to zirconolite from potassic lavas (De Hoog and Van Bergen, 2000) and lunar basalts (Rasmussen et al., 2008; Williams and Gieré, 1996). Actually, Potosí zirconolite contains the highest Y_2O_3 content described to date in terrestrial zir-

conolite (up to 11.06 wt.% Y_2O_3) and the REE contents are also high (up to 10.7 wt.% REE₂O₃), only comparable with contents in zirconolite from lunar basalts (Rasmussen et al., 2008; Williams and Gieré, 1996). The studied zirconolite contains low amounts of actinides (U, Th) strongly differing from the Finero chromitites' zirconolite (Zaccarini et al., 2004).

It is a great challenge to reconcile the different settings that this complex mineral assemblage may suggest, including SCLM, kimberlites, carbonatites, alkaline basalts, and those of lunar origin, in the suprasubduction zone ophiolite setting of Moa-Baracoa. In this context, it is essential to consider a simple integrated model that can be conceptualized in a suboceanic mantle setting as described hereafter.

5.2. Formation of the HFSE-bearing assemblage

The clinopyroxene cores crystallized at an early stage and show trace element compositions (Fig. 12c; Appendix 12) similar to those from gabbros sampled from the Mid-Atlantic Ridge (e.g., Coogan et al., 2000). This was followed by the crystallization of plagioclase (that caused depletion in Eu) and the crystallization of the clinopyroxene REE-enriched rims (Fig. 12c) and the orthopyroxene (cf. Bodinier et al., 2008; Drouin et al., 2009; Meyer et al., 1989). During the process, the evolving liquid exsolved a sulfide liquid (Fig. 4d, 11), while relict olivine from the chromitite protolith changed their composition (Fo < 90; Appendix 5), chromite recrystallized (Fig. 5; Cr# = 0.57 – 0.70, Fe³⁺# = 0.10 – 0.30, and Mg# = 0.31 – 0.53), and the HFSE-bearing mineral phases (explained hereafter) crystallized.

5.2.1. Ti-rich amphibole

The chemical composition of amphibole (Fig. 6; Appendix 5), characterized by high Ti, relatively high F, and low Cl contents indicates a magmatic origin (Coogan et al., 2001; Vanko and Stakes, 1991), though it is texturally late (coronas around plagioclase and clinopyroxene, Fig. 4a, e, and larger interstitial crystals, Fig. 4f, g). Similar compositions are found in oceanic environments, such as gabbros recovered from the Mid Atlantic Ridge (e.g., Coogan et al., 2000; 2001; Cortesogno et al., 2000; Gillis et al., 1993; Koepke et al., 2005) and in olivine-bearing gabbros from the Northern Italy ophiolites (e.g., Tribuzio et al., 1995, 2000, 2014), interpreted as the result of interaction at high temperature with percolating H_2O -rich igneous agents (e.g., Tribuzio et al., 2000). However, these differ with the compositions of amphibole from oceanic veins (compilation by Pilet et al., 2011).

Ti-rich amphibole in oceanic rocks is also typical of low-pressure (< 4 kbar) high temperature (ca. 700 - 800 °C) ocean floor metamorphism of mafic rocks, as recorded in tectonic blocks of amphibolite (former fore-arc basaltic dykes intruding residual peridotites) from serpentinite mélanges of eastern Cuba (Lázaro et al., 2016). However, amphibole formed during oceanic metamorphism is typically characterized by a high-Cl content (e.g., Currin et al., 2018; Gillis et al., 1993; Gillis and Meyer, 2001; Gillis and Thompson, 1993; Pertsev et al., 2015; Prichard and Cann, 1982; Vanko, 1986), yet the Cl content of the studied amphiboles is low (< 0.05 wt.% Cl, < 0.016 Cl per 23 O in the amphibole formula), thus dismissing a hydrothermal/metamorphic alteration at the oceanic stage. Furthermore, even if the temperature conditions of formation of Ti-rich amphibole in the studied case cannot be accurately established due to the open-system behavior during metasomatism and the lack of thermodynamic equilibrium in the studied assemblage, the Ti-Al-calcic amphibole thermobarometer of Ernst and Liu (1998) for rutile/titanite/ilmenite-saturated MORB rocks point to magmatic conditions for amphibole formation (ca. 1075 °C, 1 kbar; ca. 925 °C at 5 kbar). Fractional crystallization of a Ti-rich evolved crystallizing/reacting metasomatic melt can satisfactorily explain the enrichment in HREE of amphibole (Fig. 12b;



Fig. 10. (a) Ca-Zr-Ti (atom %) diagram of Potosí zirconolites. The reference compositional fields include data from carbonatites (Amores-Casals et al., 2019; Bulakh et al., 1999; Della Ventura et al., 2000; Hurai et al., 2017a, 2017b; Kovalchuk and Shumilova, 2014; Williams and Gieré, 1996), chromitites (Zaccarini et al., 2004), metarodingites (Stucki et al., 2001), intrusive rocks (Bellatreccia et al., 2002; Santos et al., 2015; Williams and Gieré, 1996), kimberlites (Williams and Gieré, 1996), lunar samples (Rasmussen et al., 2008; Williams and Gieré, 1996), metacarbonates (Williams and Gieré, 1996), metacarbonates (Williams and Gieré, 1996), metasomatized mantle (Della Ventura et al., 2000; Williams and Gieré, 1996), placer deposits (Williams and Gieré, 1996), potassic lavas (De Hoog and van Bergen, 2000; Carlier and Lorand, 2003, 2008), skarns (Pascal et al., 2009; Williams and Gieré, 1996), and ultramafic complexes (Azzone et al., 2009; Williams and Gieré, 1996); (b) Ca²⁺-(Th+U)⁴⁺-REE³⁺ (atom %) diagram of Potosi zirconolites. The reference compositional fields include the same data as (a); (c) Variation of Y₂O₃ (wt.%) versus HfO₂ (wt.%) in the Potosi zirconolites. The reference fields include the same data as (a); (d) Dominant coupled substitution reaction in the Potosí zirconolite. The diagram shows variation of [2Ca+Ti] versus [2(REE+Y)+(Fe+Mg)].

Coogan et al., 2000). On the other hand, the trace-element composition of studied amphibole plots in the suprasubduction environment (Fig. 13; Coltorti et al., 2007), in agreement with the geodynamic position of the studied ophiolitic chromitites in a back-arc position (Proenza et al., 1999b).

5.2.2. Mg-rich ilmenite

Ilmenite probably crystallized during or just right after the formation of amphibole as a result of Fe-Ti saturation of the melt, as indicated by the textural positions of recrystallized chromite, ilmenite, and Ti-rich amphibole. Even though Mg-rich ilmenite is scarce in basaltic rocks of the oceanic lithosphere, it has been observed in srilankite-bearing gabbroic veins that crosscut oceanic peridotites (Morishita et al., 2004).

Since ilmenite appears on the liquidus of strongly fractionated high Fe-Ti melts (Clague et al., 1981; Dick et al., 2000; Juster et al., 1989; Morishita et al., 2004 and references therein), extreme differentiation of a percolating MORB-type melt during its ascent within the uppermost mantle may be needed (c.f., Morishita et al., 2004). However, Mg-rich ilmenite cannot simply form by fractional crystallization of basaltic liquid, because Mg and Ti have opposite behaviors during this process. Following Lorand and Gregoire (2010), the interaction of an evolved (Fe-) Ti-rich melt with host Ol-bearing chromitite is necessary to explain the observed decrease in Mg in metasoma-



Fig. 11. Schematic relations between the observed mineral assemblage in the metasomatized chromitite, located at the interaction zone between the gabbroic dikes and the ordinary chromitite.

tized/recrystallized chromite and olivine and the corresponding increase in Mg in the crystallizing melt. Furthermore, preferential uptake of Zr (and Hf) by Ti-rich amphibole (e.g., Tiepolo et al., 2001) causes enrichment in Nb (and Ta) in the residual melt, explaining the positive anomalies in Nb and Ta of the Potosí ilmenites that crystallized during or shortly after precipitation of Ti-rich amphibole.

5.2.3. Apatite

An increase in P content during the magmatic/metasomatic evolution of the differentiated gabbroic melt explains crystallization of apatite in contact with, or as inclusions within, ilmenite. The Cl content in apatite is within the range of normal apatite from gabbros (Zhang et al., 2017), while the F-rich component of the Potosí apatite indicates magmatic crystallization from an igneous silicate melt (O'Reilly and Griffin, 2000), strengthening the idea that seawater-derived fluids did not participate in the metasomatic process of the Potosí chromitite. However, the trace elements in apatite are one order of magnitude depleted compared to typical MORB apatite (Mao et al., 2016) and ophiolitic Fe-gabbro apatite (Tribuzio et al., 2000), indicating a late formation of this mineral in the metasomatic process.

5.2.4. Zr minerals

5.2.4.1. Zr oxides: zirconolite, baddeleyite, and srilankite

The observed textures, with simple and multiple (e.g., baddeleyite and zirconolite or baddeleyite and srilankite; Fig. 7) Zr oxides located within and adjacent to crystal borders of ilmenite hosted within large chromite grains (up to 5 mm), suggest exsolution of Zr oxides from ilmenite (Fig. 7). Direct exchange of Zr^{4+} for Ti⁴⁺ is possible within the structure of ilmenite (Fujimaki et al., 1984; Jang and Naslund, 2003; McCallum and Charette, 1978; McKay et al., 1986; Morisset and Scoates, 2008; Nakamura et al., 1986; Taylor and McCallister, 1972). Though the amount of Zr in ilmenite is very low (up to 79 ppm; Appendix 12), we consider that the exsolution explains the tex-

tural relations of Zr oxides, as for similar baddelevite lamellae/blebs occurring within ilmenite from layered intrusions (Barkov et al., 1995; Heaman and LeCheminant, 1993; Naslund, 1987) and mafic granulite/ amphibolite (Bingen et al., 2001). A simultaneous oxidation process, as proposed by Naslund (1987) and Bingen et al. (2001), is not evident in Potosí as far as Mg-ilmenite contains a limited amount of trivalent Fe and lacks hematite exsolutions. A second generation of sheet-like baddelevite, together with Mg-Al silicates, crystallized after zircon in Potosí (Fig. 9c). Similar textures were reported by Morisset and Scoates (2008), where baddeleyite forms along sealed fractures within zircon and at zircon grain rims. Zirconolite and srilankite likely formed by an exsolution mechanism similar to baddeleyite however it depended on the availability of Ca and REE, which was provided by the evolved melt and/or apatite that also formed in contact with ilmenite. Thus, zirconolite would concentrate incompatible REE+Y of the late magmatic liquids/fluids.

5.2.4.2. Zircon

For many Fe-Ti oxide gabbros, zircon formed from volumetrically small, late-stage, fractionated, Zr-enriched interstitial melt pockets (Scoates and Chamberlain, 1995; Grimes et al., 2009; Beckman et al., 2017; but see also Pietranik et al., 2017, and references therein, for formation of zircon slightly above the wet basaltic solidus after fluid-fluxed partial melting of oceanic gabbros). Zircon occurs in Cuban ophiolitic gabbros and chromitites as magmatic (neoformed) and xenocrystic grains inherited from subducted sedimentary rocks (Proenza et al., 2018; Rojas-Agramonte et al., 2016). However, the studied zircon grains in Potosí are clearly not magmatic s.s. nor xenocrystic, but rather formed *in situ* during metasomatic interaction of chromitite with gabbroic melts, as indicated by textural relations.

Zircon coronas around ilmenite suggest high temperature Zr diffusion in ilmenite (exsolution) and subsequent reaction along grain boundaries under supersolidus to subsolidus conditions following ilmenite crystallization (Morisset and Scoates, 2008). As indicated above. Zr is highly compatible within the ilmenite structure and this mineral is usually considered as one potential source of Zr for zircon formation (Bea et al., 2006; Charlier et al., 2007; Morisset and Scoates, 2008; Sláma et al., 2007). The latter authors demonstrated that a reasonable concentration of Zr in ilmenite (345 - 568 ppm Zr initial concentration) accounts for the observed amount of zircon formed at the rims of ilmenite in mafic plutonic rocks of Proterozoic anorthosite suites. Furthermore, the studied zircon crystals have very low U and Pb contents, consistent with derivation from ilmenite (Morisset and Scoates, 2008; Zack and Brumm, 1998). Formation of zircon after baddelevite exsolutions in the presence of silica, as proposed by Beckman et al. (2017), cannot be considered here because the observed relationship involves secondary growth of baddeleyite after replacement of granular zircon grains, and the presence of ilmenite interstitial to the granular zircon grains of the coronas is consistent with direct formation of zircon after ilmenite. In this process, silica would be provided by the adjacent silicates and/or the residual melt.

A metamorphic origin for zircon after high-pressure and/or temperature reactions (e.g., Austrheim et al., 2008; Beckman et al., 2014, 2017; Bingen et al., 2001; Kovaleva and Klötzli, 2017; Söderlund et al., 2004) is not considered here due to the lack of evidence for solid-state metamorphic transformation in the studied rocks. Likewise, a hydrothermal origin, similar to the zircons hosted in sealed fractures within chromitites introduced via a network of high-T fluids emanating from gabbroic dikes (Kapsiotis et al., 2016), is also discarded because of the low HfO₂ content of the Potosi zircons (< 1.54 wt.%; Appendix 11) and the HREE enrichment relative to LREE (Fig. 12g). Hydrothermal zircon usually has high HfO₂ content (3.4 to 4.9 wt.%) (Hoskin and Schaltegger, 2003), high U (550 – 13,000 ppm) and



Fig. 12. (a) Chromite trace elements normalized to accessory MORB chromite (Pagé and Barnes, 2009). The regional Moa-Baracoa Al-rich chromite data is from Colás et al. (2014); (b) Amphibole REE normalized to chondrite C1 (Sun and McDonough, 1989). Amphibole in ophiolitic Fe-gabbros is from Tribuzio et al. (2000) and amphibole in oceanic lithosphere veins is a compilation by Pilet et al. (2011); (c) Clinopyroxene REE normalized to chondrite C1. Note the different patterns in the core and rim of clinopyroxene. Clinopyroxene in oceanic gabbros is from Coogan et al. (2000) and clinopyroxene in equilibrium with MARK MORB is from Reynolds and Langmuir (1997); (d) Orthopyroxene REE normalized to chondrite C1. Orthopyroxene in agbbronorites is from Secchiari et al. (2018); (e) Ilmenite trace elements normalized to chondrite C1. Mg-rich ilmenites in amphibole-rich veins from Lherz are from Lorand and Gregoire (2010); (f) Apatite REE normalized to chondrite C1. Apatite in ophiolitic Fe-gabbros is from Tribuzio et al. (2000) and apatite in MORB is from Mao et al. (2016); (g) Zircon REE normalized to chondrite C1. Magmatic oceanic crust zircons are from Grimes et al. (2009); and (h) Zirconolite REE normalized to chondrite C1. Lunar zirconolite is from Williams and Gieré (1996) and Rasmussen et al. (2008).



Fig. 13. Classification diagrams of amphiboles: (a) Zr/Nb vs Ti/Nb; and (b) Zr/Nb vs Ti/ Zr. Fields for intraplate and suprasubduction amphiboles are from Coltorti et al. (2007).

Th (450 – 6,000 ppm) (Belousova et al., 2002), and relatively flat chondrite-normalized REE patterns (Hoskin, 2005).

5.3. Fractional crystallization and interaction of evolved melts with chromitite

In the Moa-Baracoa ophiolite, Marchesi et al. (2007) demonstrated that the primitive lower crust layered gabbros and pegmatitic gabbros have similar Sr and Nd isotopic signatures and derived from a common isotopic DMM source, which is similar to that of the Potosí pegmatitic gabbros (in POT-1 sample; ⁸⁷Sr/⁸⁶Sr = 0.702581 and 143 Nd/ 144 Nd = 0.513128; unpublished data). It is then conceivable that the Potosí pegmatitic gabbro dikes formed from derivative liquids after the crystal fractionation of primitive MORB similar to those that built the Moa-Baracoa lower oceanic crust and the more primitive olivine gabbro dikes intruding Potosí chromitites. Like oceanic Fe-Ti gabbros in slow-spreading mid-ocean ridges (Chen et al., 2019; Dick et al., 2000), Potosí gabbros have an assemblage composed of clinopyroxene and plagioclase - with multiple oxide inclusions - that is distinctive of gabbroic rocks derived from MOR-like basaltic melts (with BABB-like trace element signature, Fig. 3). Under the quartz-fayalite-magnetite oxygen buffer condition, primitive MORB crystallizes olivine, clinopyroxene, plagioclase and, upon further cooling below c. 1100 °C, Fe-Ti oxides forming oxide gabbros (e.g. Chen et al., 2019), in agreement with the mineral assemblage observed in Potosí gabbro dikes. The low anorthite content of plagioclase (An₅₁₋₅₆; Appendix 5) and a rather high Mg# (85 - 90 mol.%) of Potosí pegmatitic gabbros are consistent with crystallization from a rather anhydrous (< 0.5 wt.%) evolved MORB melt for a liquidus temperature of c. 1100 °C (Berndt et al., 2004; Feig et al., 2006). However, their high contents in REE and incompatible trace elements (Fig. 12, Appendix 12), their unfractionated trace element patterns and the strong enrichments in trace elements in the rims of high-Mg# clinopyroxene cannot solely be accounted by crystal fractionation. These geochemical features are symptomatic of post cumulus fractionation of melts in a solidifying crystal-melt mush and entrapment of residual evolved melts upon cooling (e.g. Bédard, 1994; Bodinier et al., 1986; Borghini and Rampone, 2007; Garrido et al., 2006; Langmuir, 1989). This trapped melt mechanism is well documented in continental igneous layered intrusions (Cawthorn and Tegner, 2017; Ibanez-Mejia and Tissot, 2019). In this process, the residual melts are progressively and effectively isolated from early cumulus minerals due to physical or kinetic barriers and evolve via intercumulus fractional crystallization upon cooling. Borghini and Rampone (2007) successfully modeled trace element enrichment in MORB-type primitive olivine-rich cumulates from the Erro-Tobio mantle peridotite by small-scale migration of residual evolved melt by crystal fractionation and entrapment of low fractions (5 - 10%). The resulting melts are enriched in REE, Ti and Zr accounting for the occurrence of accessory interstitial Ti-amphibole and Fe-Ti oxide cumulates without requiring an exotic REE- and HFSE-rich melt (Borghini and Rampone, 2007).

Proenza et al. (2001) proposed that the intrusion of gabbroic dikes into chromitites triggered the crystallization of sulfide-rich chromite ores. Here, we propose a two-stage model for the genesis of metasomatic REE- and HFSE-rich melt in Potosí chromitites. In a first stage, crystallization of evolved MORB in a crystallizing mush produced via intercumulus fractional crystallization water-rich residual melts strongly enriched in incompatible elements. In a second stage, batches of these evolved melts escaped the solidifying mush and extensively reacted and metasomatized the surrounding chromitites, crystallizing HFSE- and REE-bearing minerals and Fe-Cu-Ni sulfides. We envisage that this process occurred in a solidifying mush in dike intrusions -in a nearly closed system or with low rate of magma supply- at near solidus conditions in the conductive oceanic mantle lithosphere, allowing melts to stagnate, slowly cool and differentiate, and the residual melts to leak and react with the surrounding chromitites and peridotites (e.g. Morishita et al., 2004).

During crystallization in the crystal mush, residual melt fractions produced via an inter-cumulus fractional crystallization were progressively enriched in FeO, TiO₂, HFSE, H₂O, promoting crystallization of HFSE- and REE-bearing minerals and, upon sulfide liquid immiscibility, the Fe-Cu-Ni sulfide assemblage described by Proenza et al. (2001). To yield residual melts with enough H₂O (2-6 wt.%) to stabilize amphibole after a primitive MORB composition (0.10-0.20 wt.% of H₂O) requires over 90% fractional crystallization driven by anhydrous mineral fractionation up to near solidus conditions (Gillis and Meyer, 2001; Tribuzio et al., 2000). This substantial degree of fractionation generated via intercumulus fractional crystallization generated near-solidus water-rich buoyant silicic melt fractions that were strongly enriched in Zr and other HFSE until the onset of zircon and baddelevite saturation (<900 °C for melts with $H_2O > 2$ wt.%; Ibanez-Mejia and Tissot, 2019). These residual melts percolated the crystallizing mush and were progressively trapped leading to the strong whole-rock trace element enrichments of the Potosí pegmatitic gabbro dikes.

Locally, batches of these HFSE-enriched residual melts escaped the crystallizing mush and tapped into the surrounding chromitites. These water-rich siliceous residual melts were in strong disequilibrium with the surrounding chromitites, leading to extensive reaction with chromitites and metasomatism, and to the rapid consumption of the melts via melt-consuming melt-rock reactions. The substantial compositional variations in the studied minerals, especially in ilmenite, and the heterogeneous distribution of the accessory Zr-bearing minerals indicates strong local heterogeneities in SiO₂, MgO, and HFSE, which is compatible with a kinetically-inhibited restricted system that evolved by rapid melt-rock reactions at decreasing melt mass. During this stage, Zr-bearing minerals (baddeleyite, zirconolite, srilankite) may have formed by exsolution after Zr-rich ilmenite, forming zircon coronas and likely involving residual silica-rich melt/ fluids at temperatures below the zircon saturation. These HFSE- rich residual melt that formed the particular mineral assemblage of the Potosí metasomatized chromitites derives from a BABB-like melt (MORB-like melt) in a suprasubduction environment (Marchesi et al., 2006, 2007; Proenza et al., 2001).

The U/Pb age of baddeleyite is indicative of the age of the metasomatic event (134.4 ± 14 Ma) that was coeval with the oceanic stage in back-arc suprasubduction settings. Similar ages were reported in igneous zircon grains from a Fe-Ti-rich gabbro intruding the mantle peridotite section of the Moa-Baracoa massif (124 ± 1 Ma; Rojas-Agramonte et al., 2016) and two zircons recovered from the Potosi ordinary chromitite (99 ± 21 Ma and 118 ± 8 Ma, Proenza et al., 2018). The metasomatic event is hence consistent with magmatism related to ophiolite construction, generally considered to have taken place at 90 - 125 Ma (Iturralde-Vinent, 1996; Iturralde-Vinent et al., 2006).

6. Concluding remarks

The interaction between chromitites and gabbroic intrusions in Potosí (eastern Cuba ophiolite) triggered the formation of a particular HFSE-mineral assemblage including Ti-rich amphibole, Mg-rich ilmenite with inclusions of Zr oxides, such as baddeleyite, zirconolite and srilankite, and zircon coronas surrounding ilmenite grains. These mineral phases related to the metasomatized chromitite are also REE-rich. We propose a two-stage model for the formation of the HFSE- and REE-bearing minerals in the Potosí metasomatized chromitites: (1) crystallization of evolved MORB-like melt (BABB-like melt in a suprasubduction environment) in a crystallizing mush, producing water-rich residual melts enriched in incompatible elements via intercumulus fractional crystallization; (2) reaction of batches of the evolved residual melts that escaped the solidifying mush, with the surrounding chromitites, giving place to the observed metasomatism and the formation of HFSE- and REE-bearing minerals and Fe-Cu-Ni sulfides. This process likely occurred near solidus conditions in a nearly closed system (or with a low rate magma supply), which allowed the melts to stagnate, slowly cool, differentiate, and react with the surrounding chromitites. The age of the metasomatic event is 134.4 ± 14 Ma, coeval with the magmatism related to ophiolite construction in eastern Cuba.

Uncited references

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Declaration of Competing Interest

The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

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Appendix A. Supplementary data

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